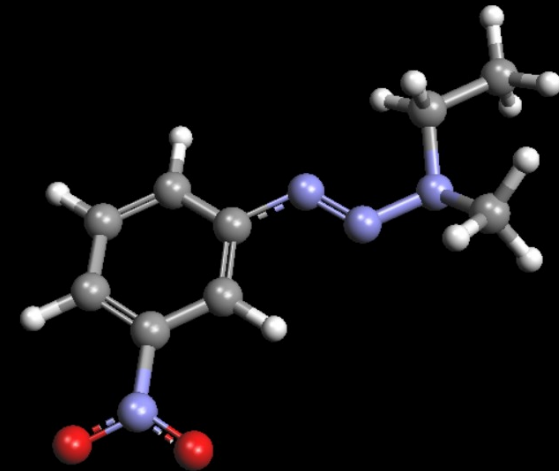




Innovative optical elements for astronomy: from the molecule to the “on sky” device

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INAF – Astronomical Observatory of Abruzzo, Colloquia
April 13, 2023

The beginning...more than 400 years ago



August 1609

Galileo Galilei showing the Doge of Venice how to use the telescope



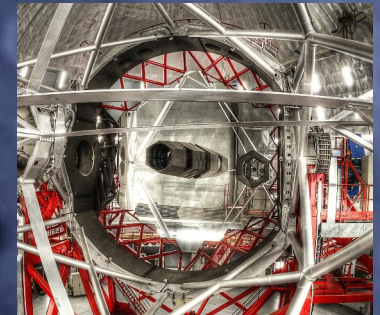
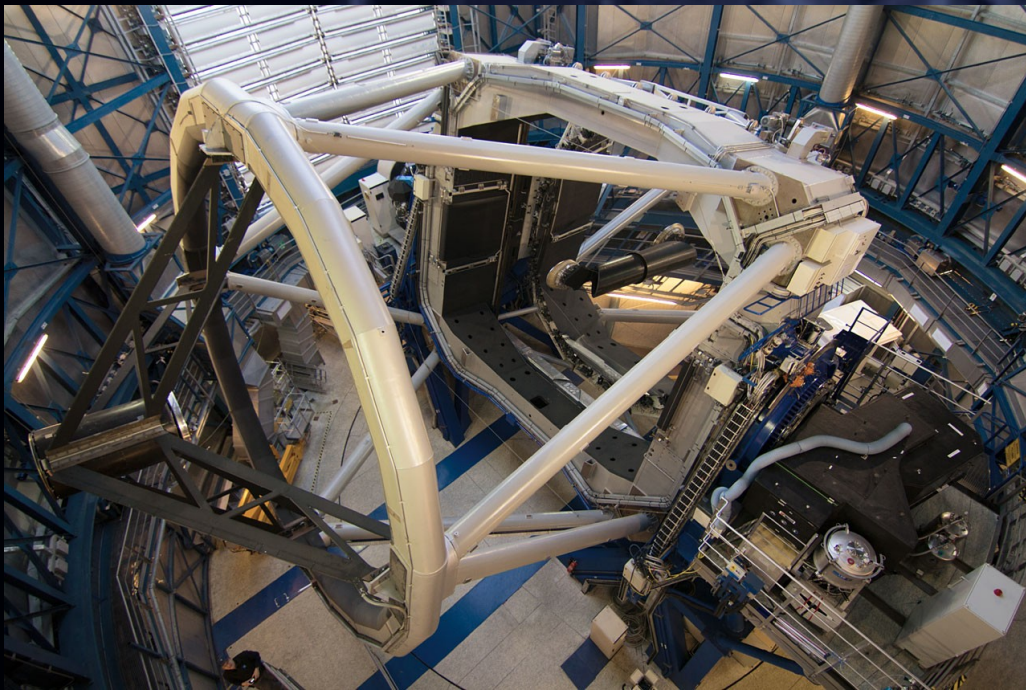
Ground based Telescopes: present

VLT: four Unit Telescopes main mirrors of 8.2m diameter

Since 2000

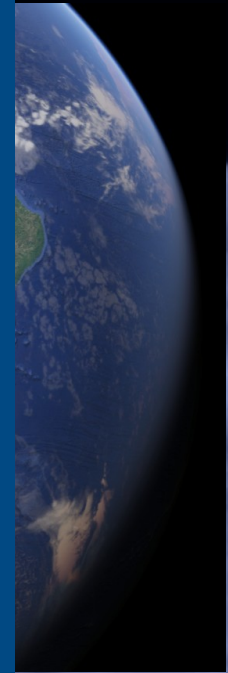
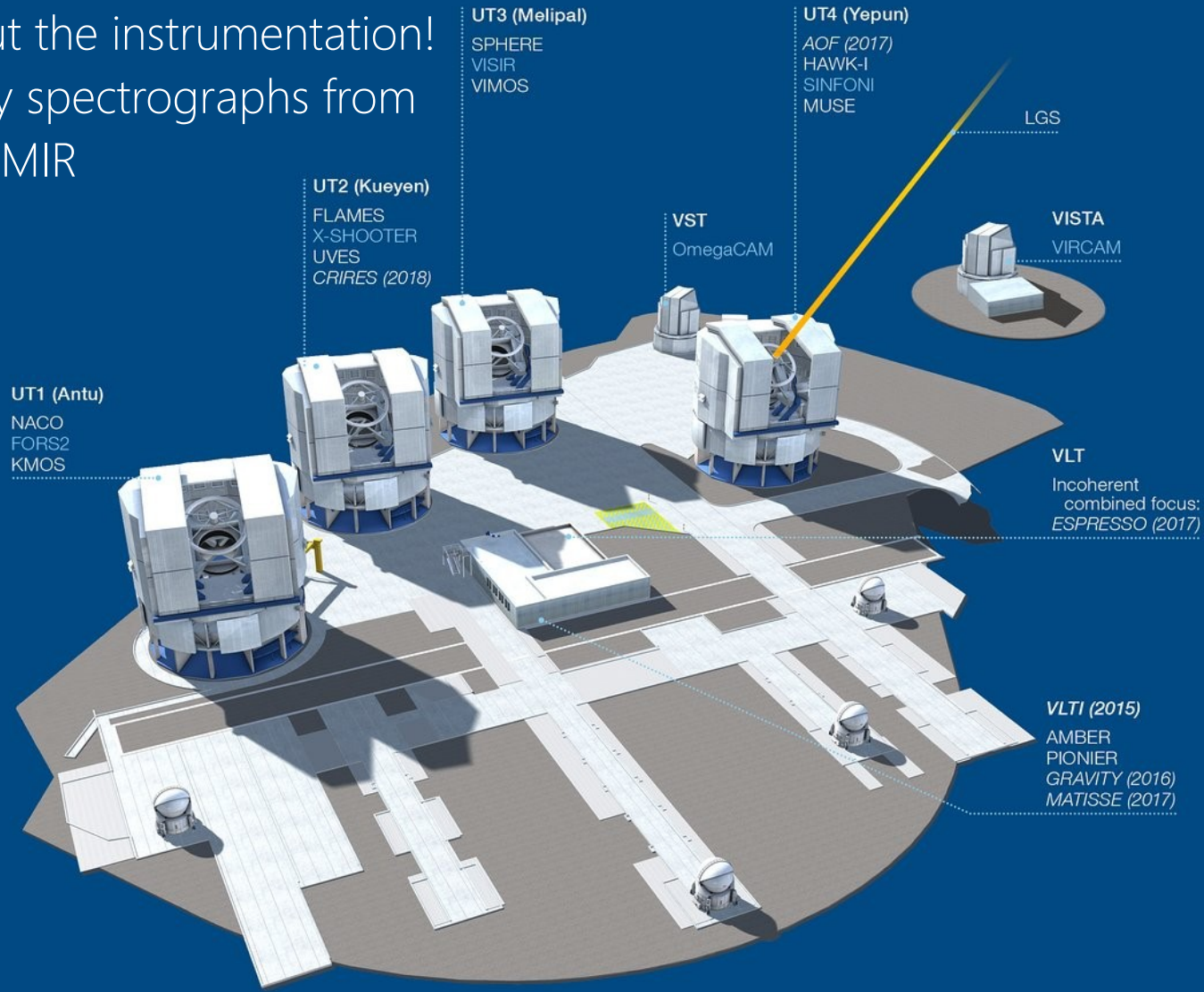


Nowadays, there are many 8-10 meter class telescopes both with monolithic and segmented mirrors.

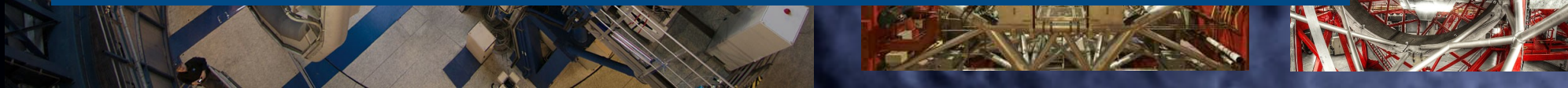
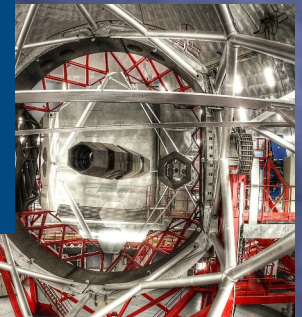


Ground based Telescopes: present

The telescope is useless without the instrumentation!
Usually spectrographs from UV to MIR

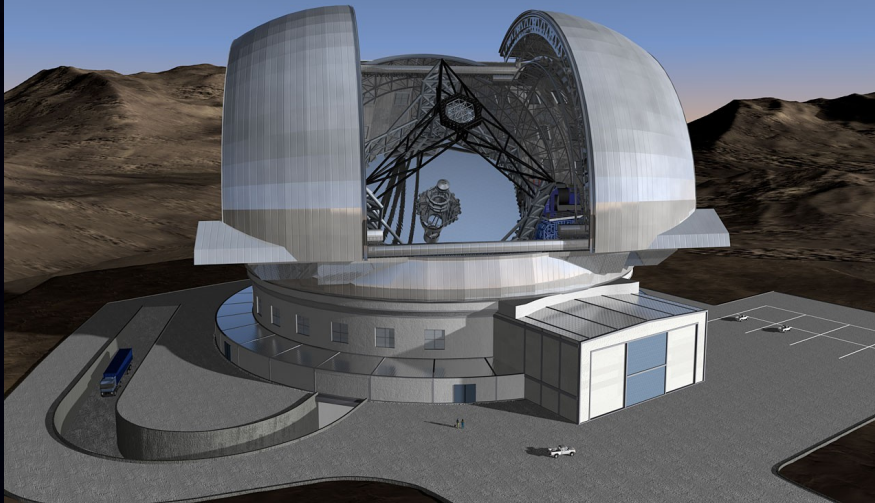


8-10 meter
monolithic



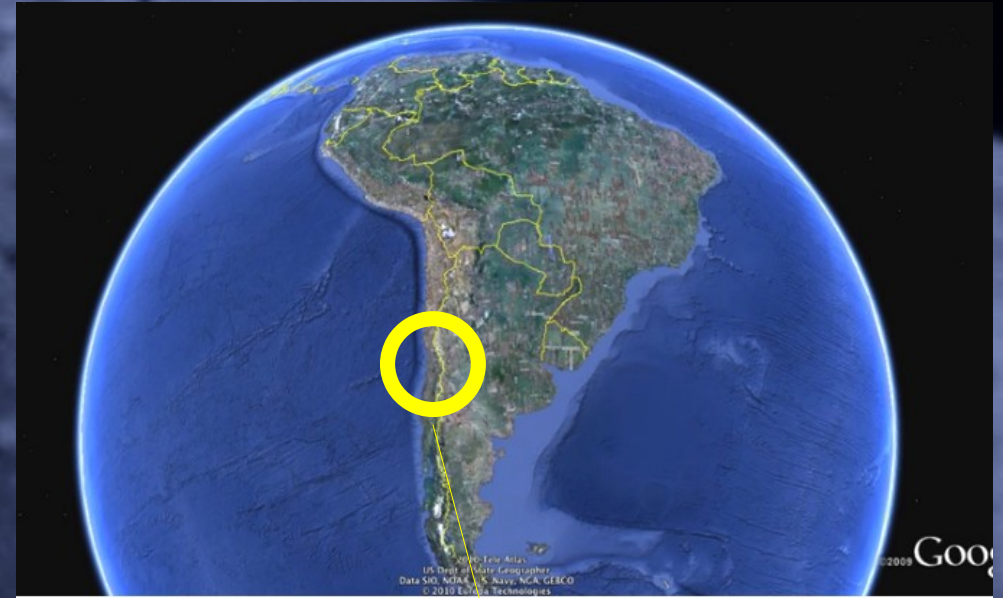
Ground based Telescope: future

First light 2028



ELT: 39 m diameter primary mirror consisting of 798 segments, each 1.4 m wide, but only 50 mm thick.

Similar ELTs in the US (GMT, TMT).

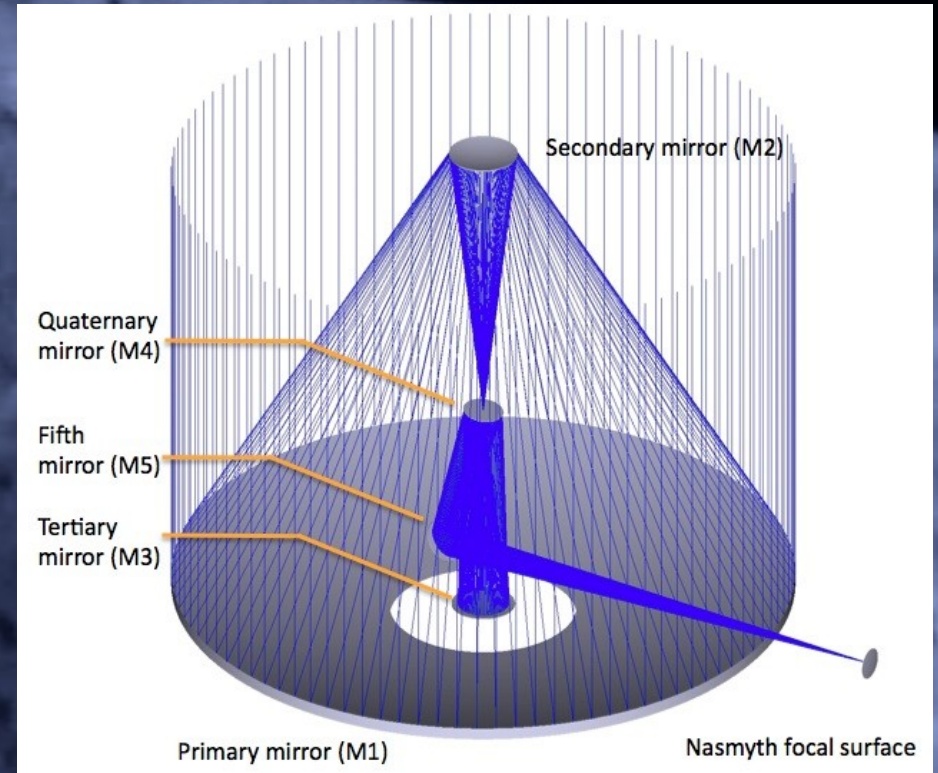


Cerro Armazones in Chile

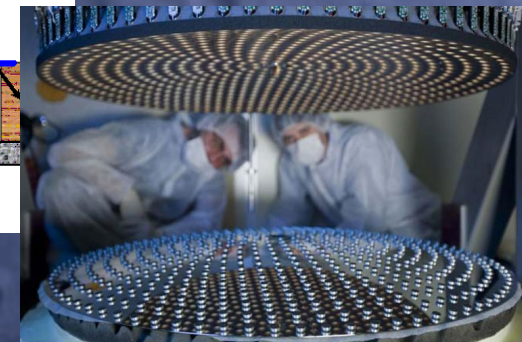
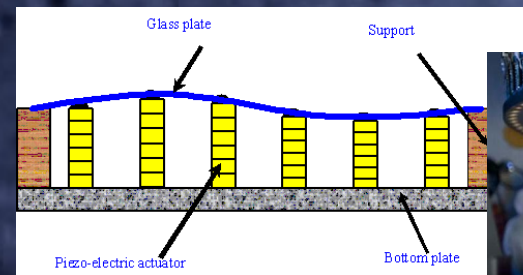


Ground based Telescope: future

First light 2028



5 mirrors, two are adaptive, M4 and M5

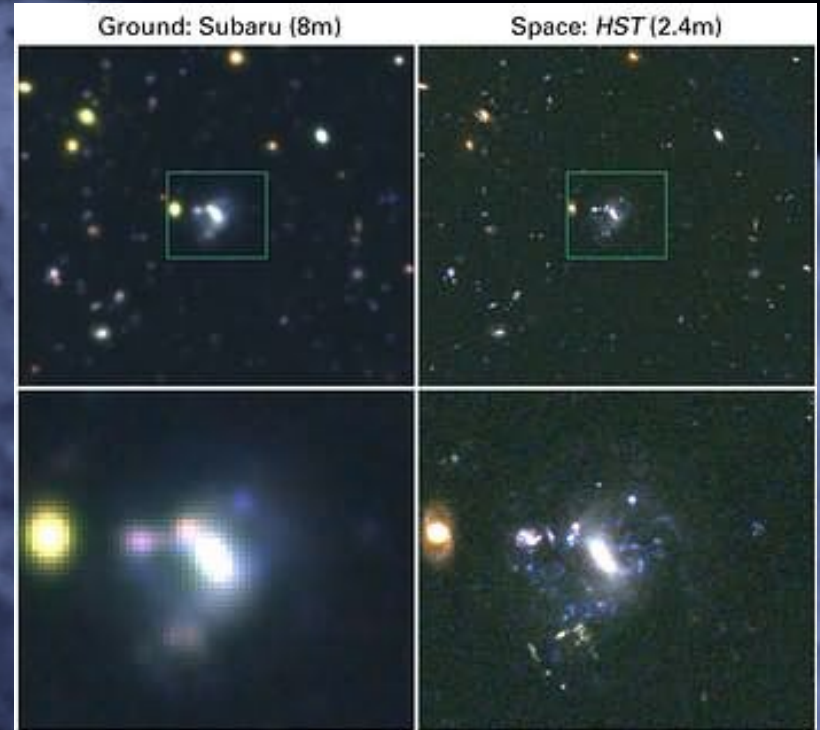


Space Telescopes: past

Hubble space telescope: 2.4 m in diameter



Since 1990



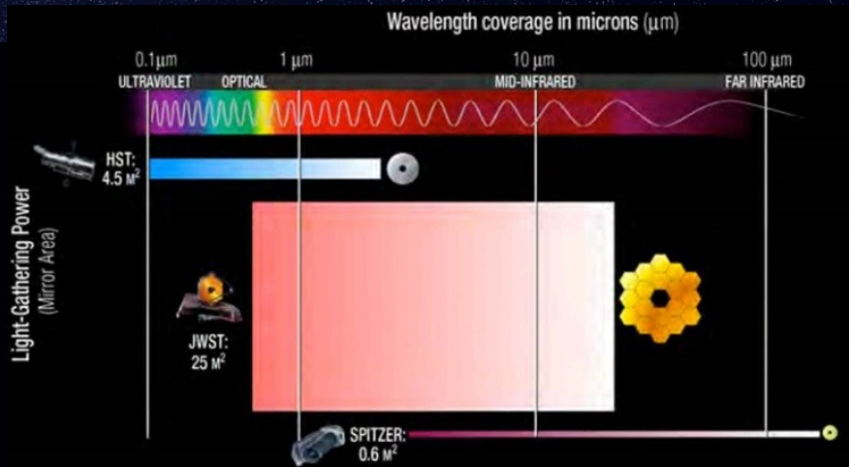
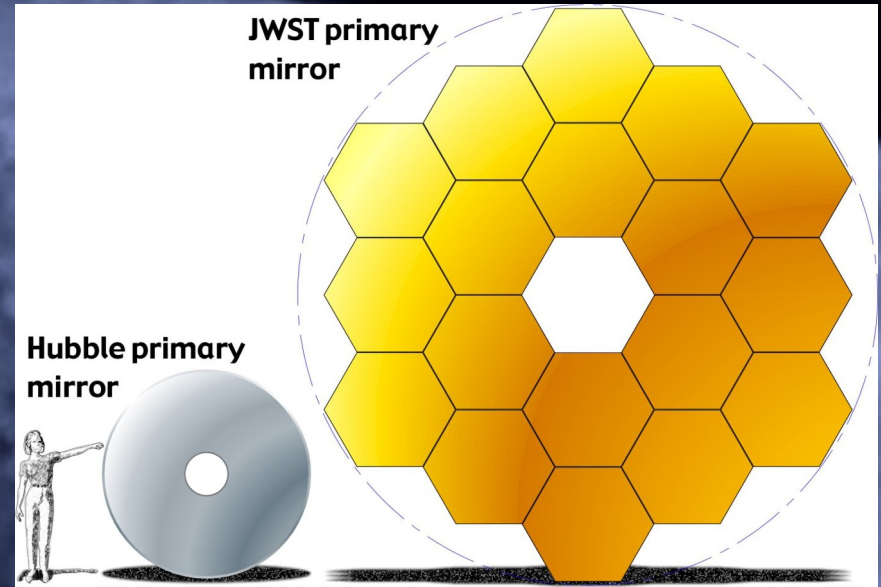
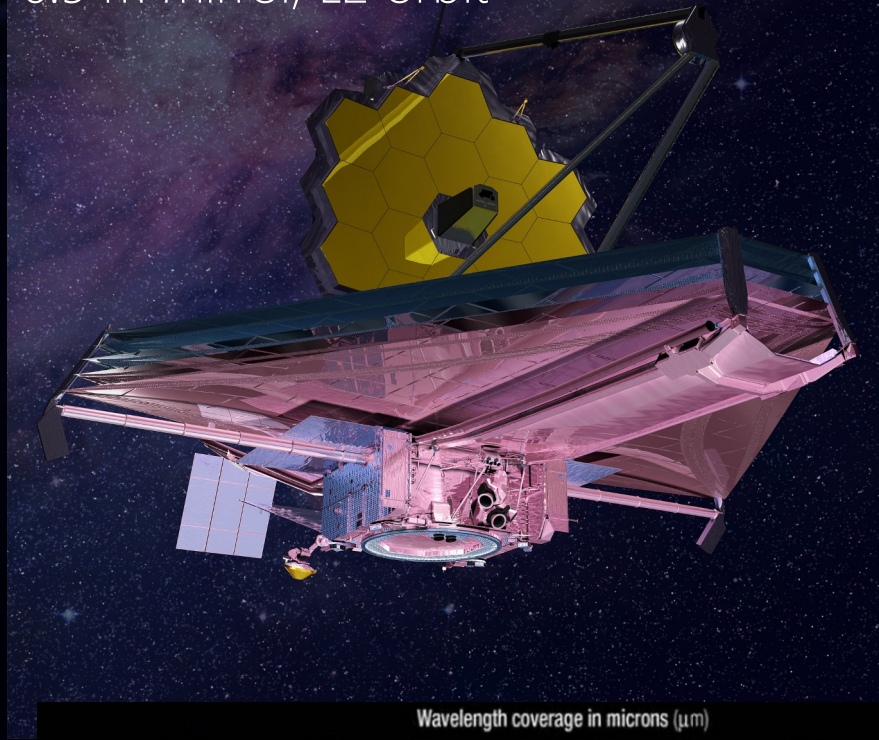
Space telescopes are not affected by the atmosphere, both turbulences and absorptions.

Hubble has four main instruments to observe in the near ultraviolet, visible, and near infrared spectra.

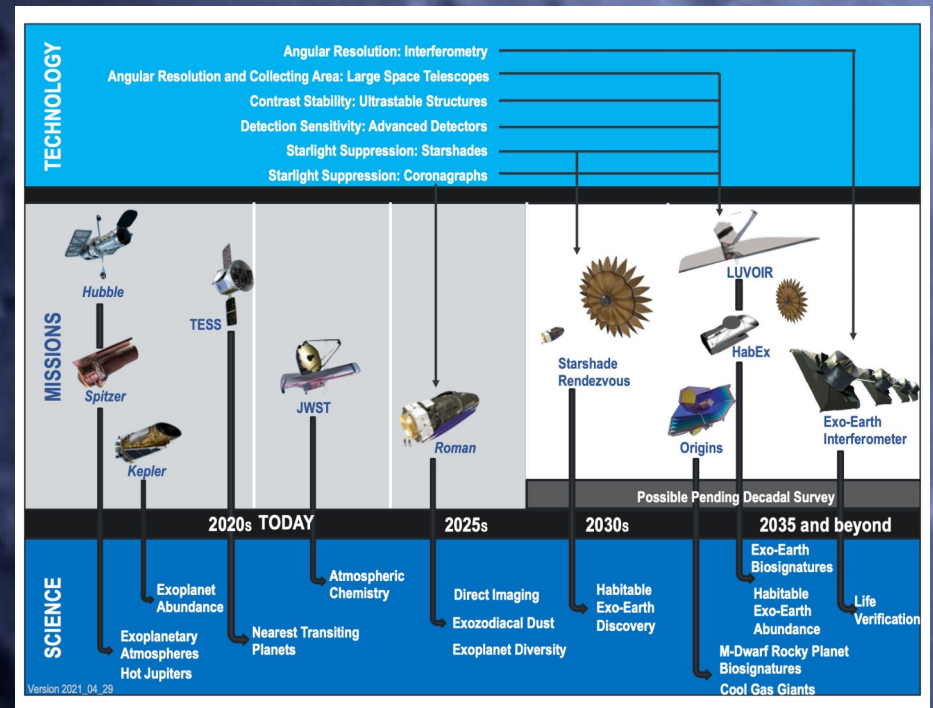
Space Telescopes: present

Launched Christmas 2021

James Webb Space Telescope:
6.5 m mirror, L2 orbit



https://youtu.be/RzGLKQ7_KZQ



Try to make stuff simpler...

Both telescopes and the instrumentation are featuring a strong increase in their complexity, due to: i) the larger sizes; ii) the demanding performances and accuracy required by the new exciting scientific cases to be faced.

The development of innovative materials and processes is a key strategy to reduce the complexity while keeping or improving the performances.

In other words...

The device complexity is "simplified" by the material.

Try to make stuff simpler...

Both telescopes and the instrumentation are featuring a strong increase in their complexity, due to: i) the larger sizes; ii) the demanding performances and accuracy required by the new exciting scientific cases to be faced.

The development of innovative materials and processes is a key strategy to reduce the complexity while keeping or improving the performances.

***From the requirements to the devices
through the developing of suitable (smart) materials
=
Multidisciplinary approach***

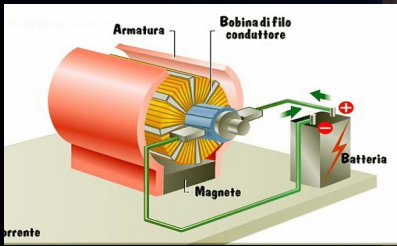
Complexity example...automatic blinds

- IDEA: implement an automatic electric blind



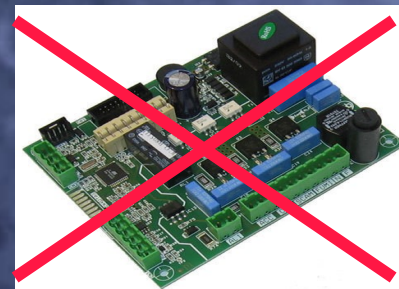
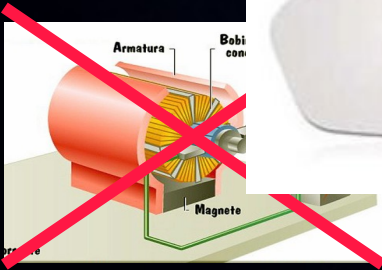
Complexity example...automatic blinds

- IDEA: implement an automatic electric blind



Complexity example...automatic blinds

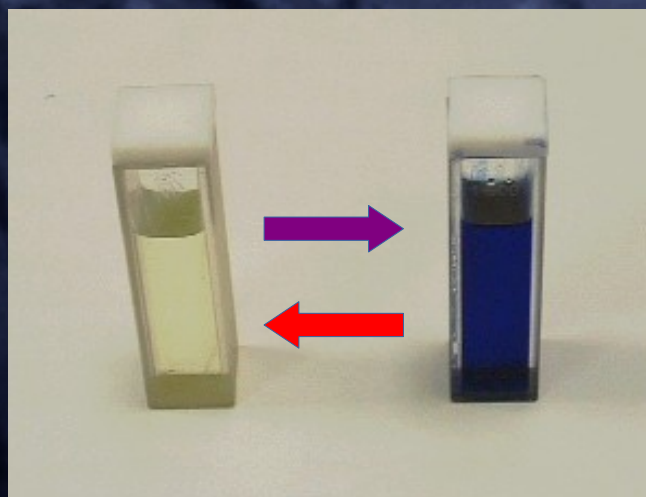
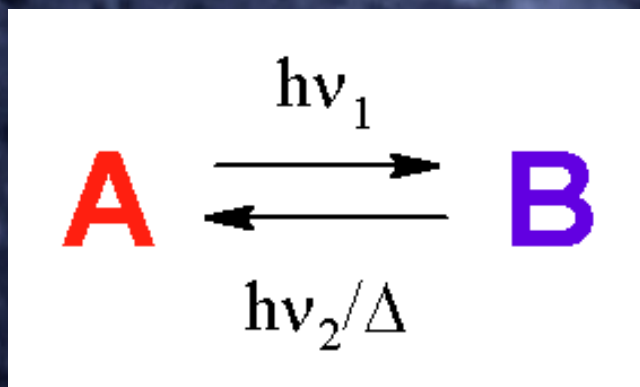
- IDEA: implement an automatic electric blind



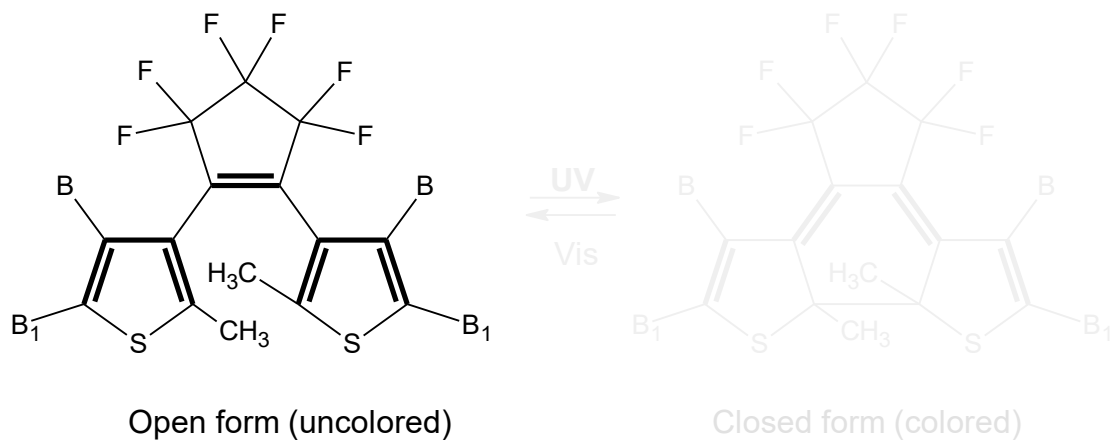
Photochromic materials

IUPAC definition:

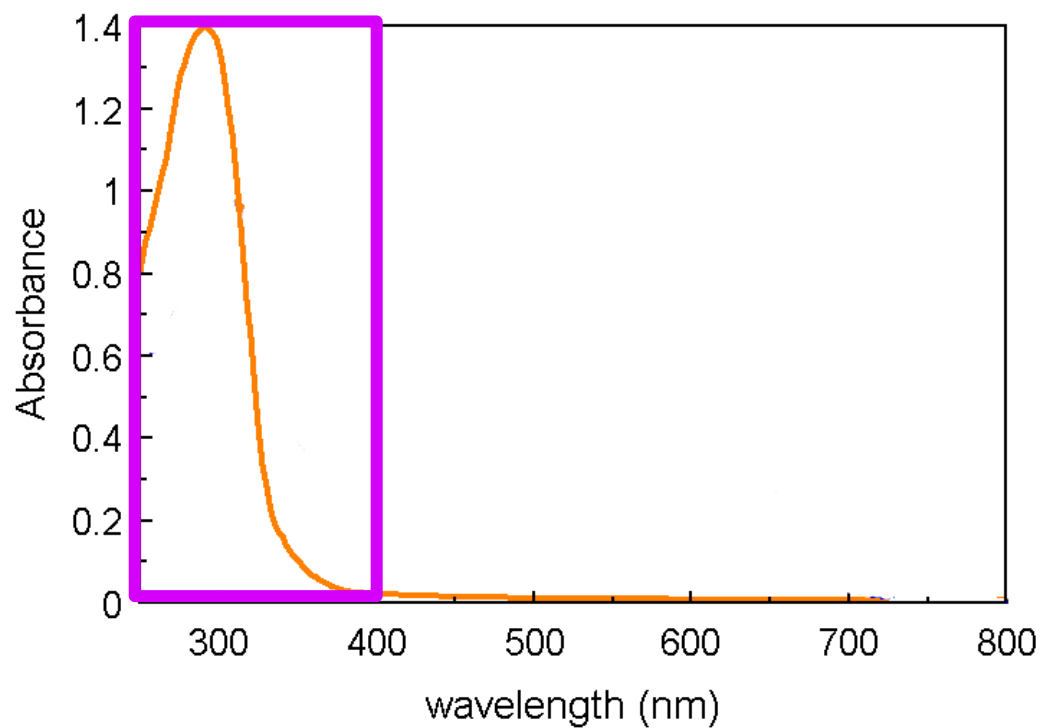
"Photochromism is a reversible transformation of a chemical species induced in one or both directions by absorption of electromagnetic radiation between two forms, A and B, having different absorption spectra"



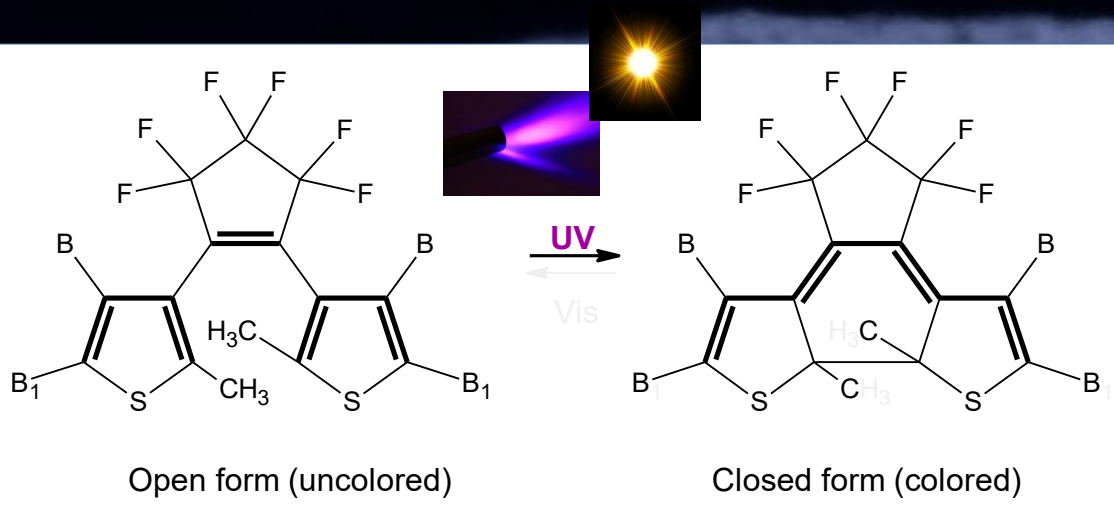
Diarylethenes



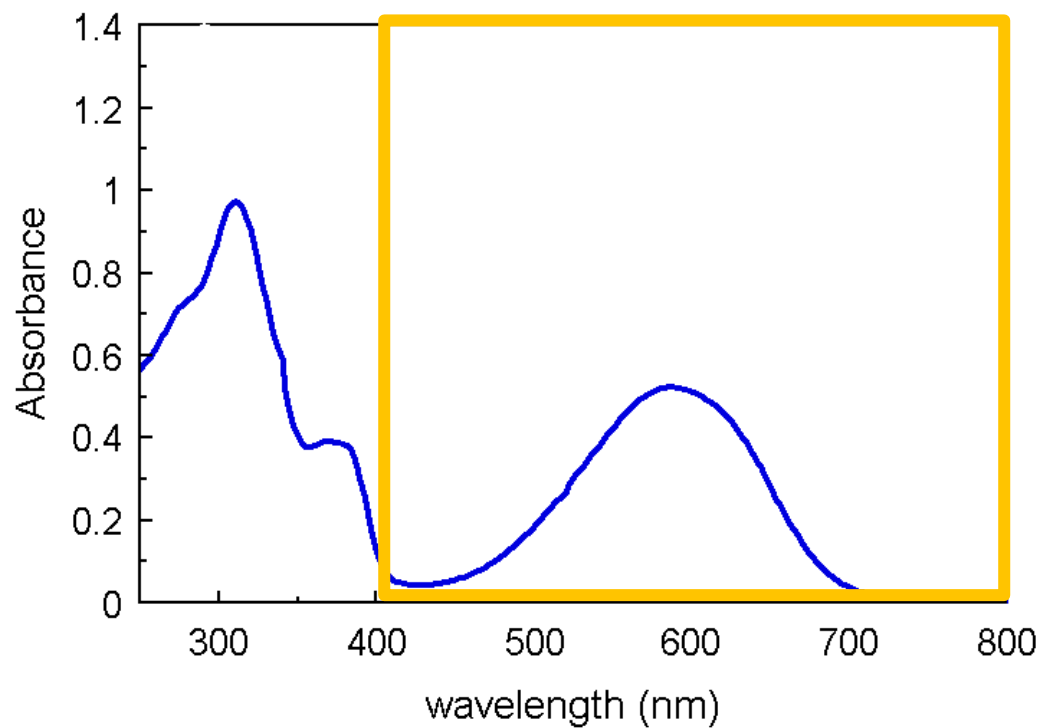
In this form, the absorption is only in the visible: uncolored material



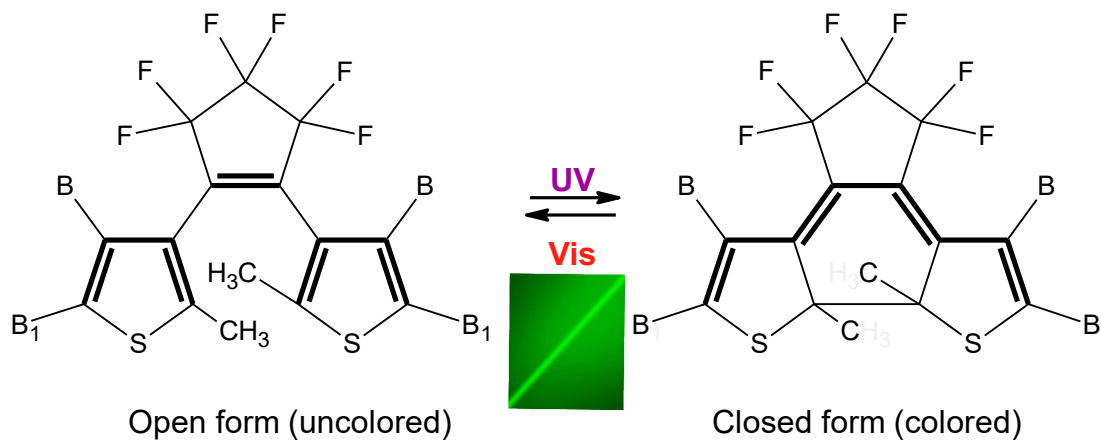
Diarylethenes



UV light
Cyclization reaction
Increase of π conjugation

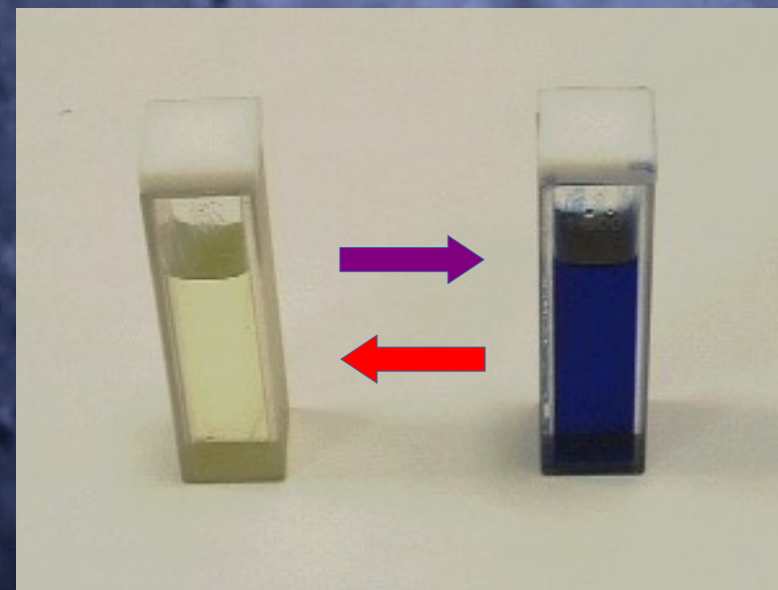
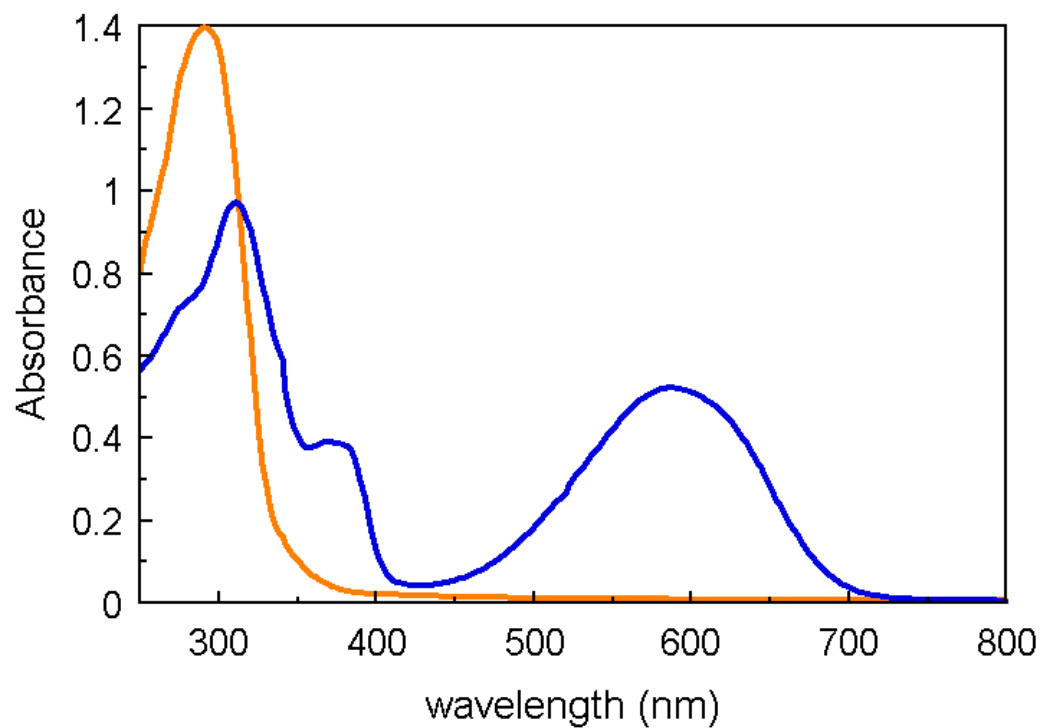


Diarylethenes

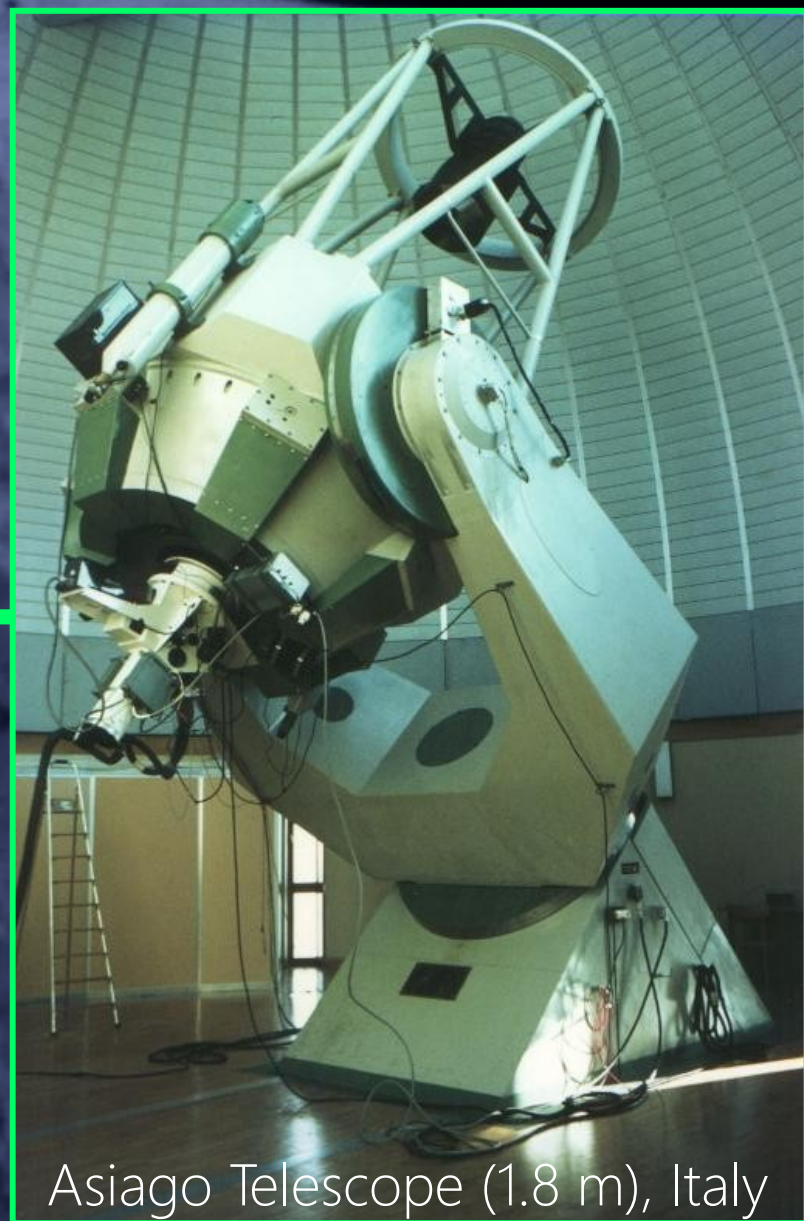
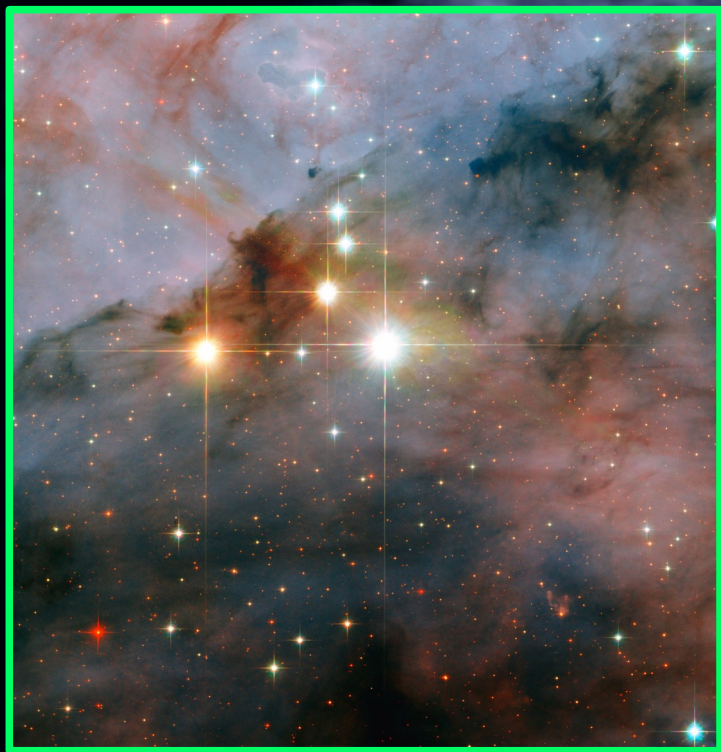


Visible light
Restoring the uncolored form
The isomer is **thermally stable**.

This is reversible!



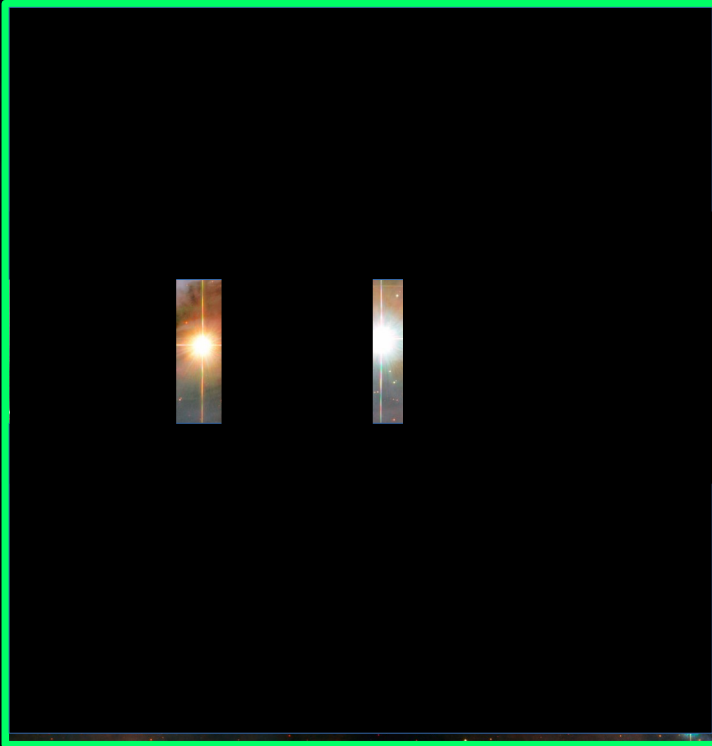
Rewritable Focal Plane Mask



Asiago Telescope (1.8 m), Italy

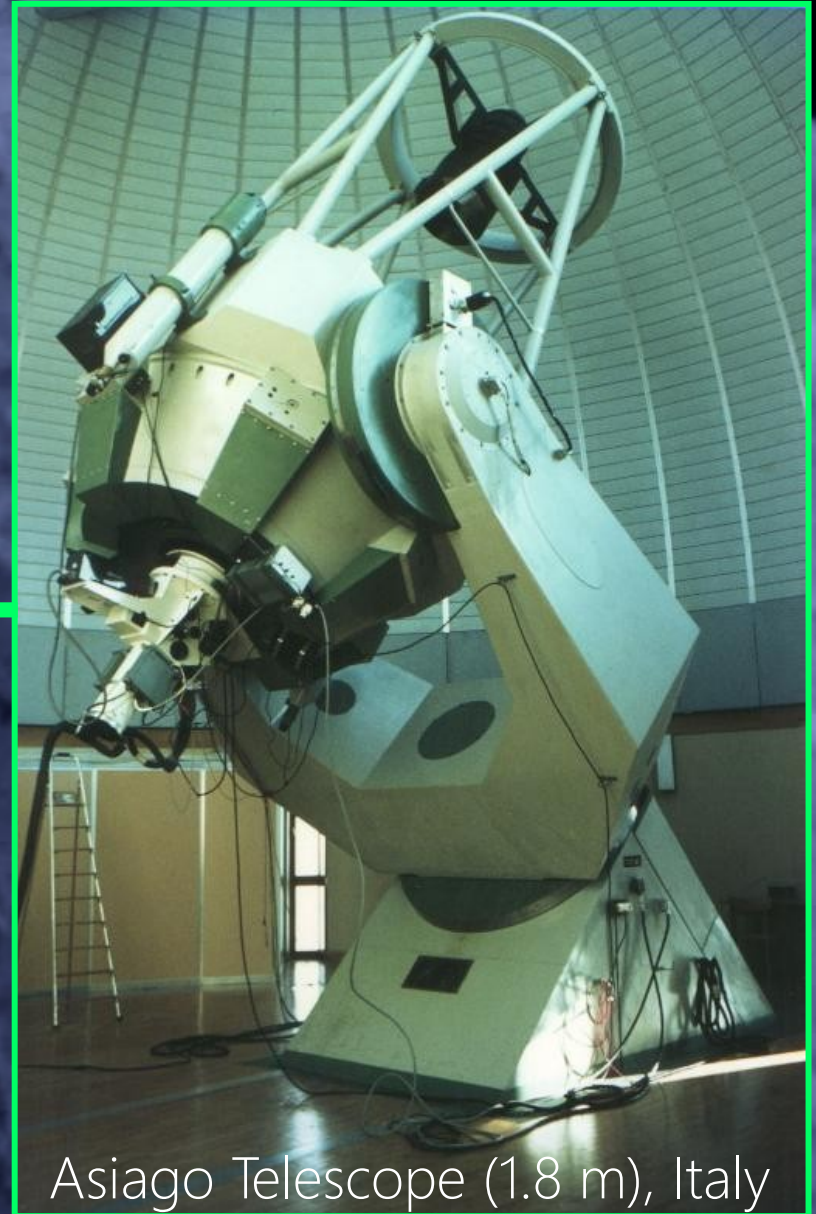
Target sky objects must be selected to record their spectra avoiding the sky contamination.
(Multi Object Spectroscopy, MOS)

Rewritable Focal Plane Mask



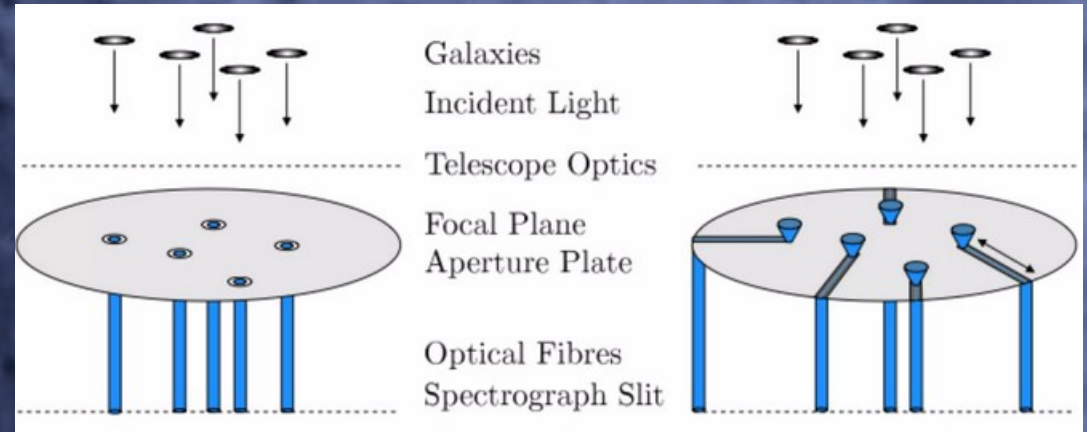
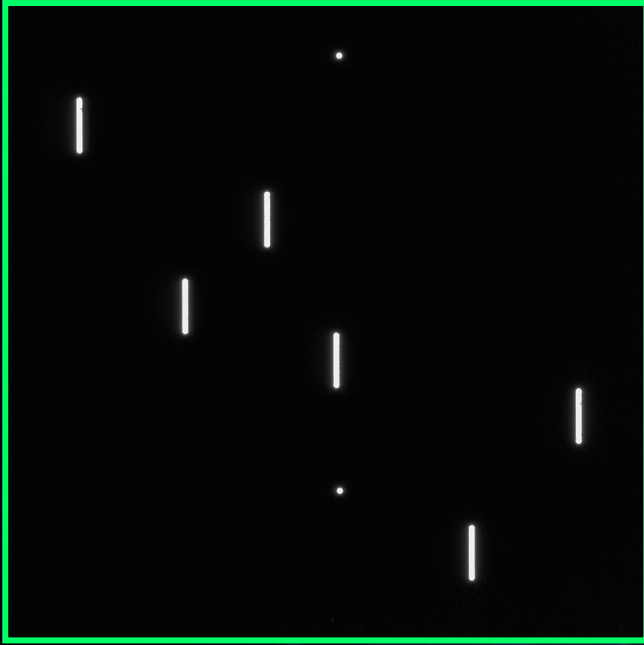
Target sky objects must be selected to record their spectra avoiding the sky contamination.
(Multi Object Spectroscopy, MOS)

FOCAL PLANE MASKS



Asiago Telescope (1.8 m), Italy

Rewritable Focal Plane Mask



Traditional focal plane mask:

- Metal sheet with cut slits
- well established technology
- highest contrast
- disposable

Optical fibers

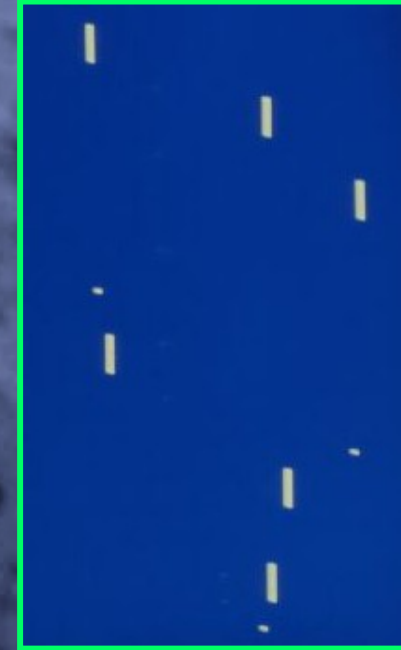
- Different approach, a huge modification of the instrument;
- Decouple the focal plane and the spectrograph;
- More complex approach.

Rewritable Focal Plane Mask



Traditional focal plane mask:

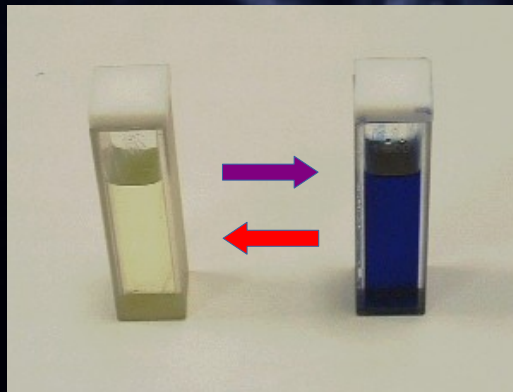
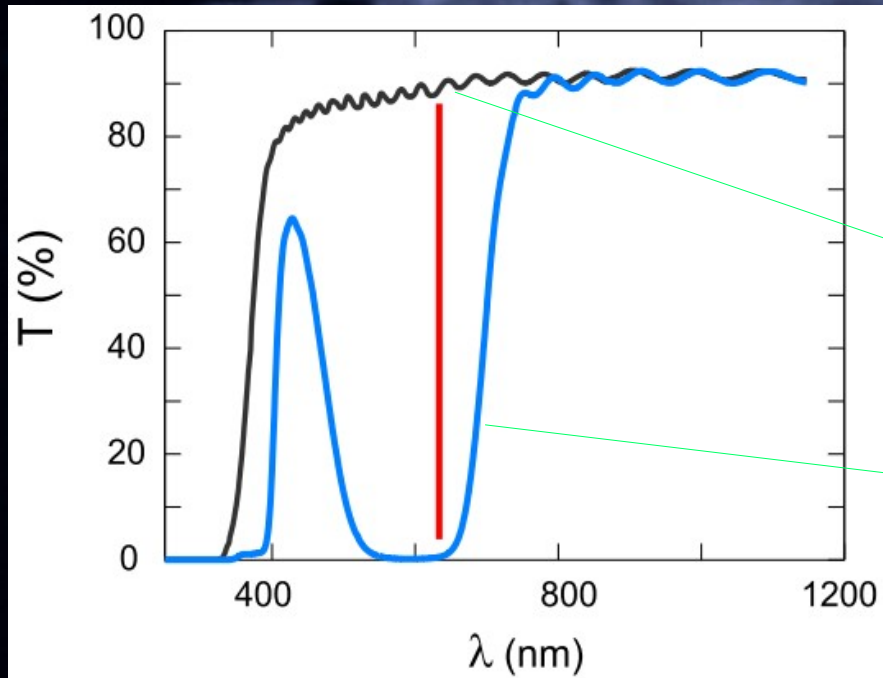
- Metal sheet with cut slits
- well established technology
- highest contrast
- disposable



Photochromic focal plane mask:

- Reusable
- Easy to use also for complex slits

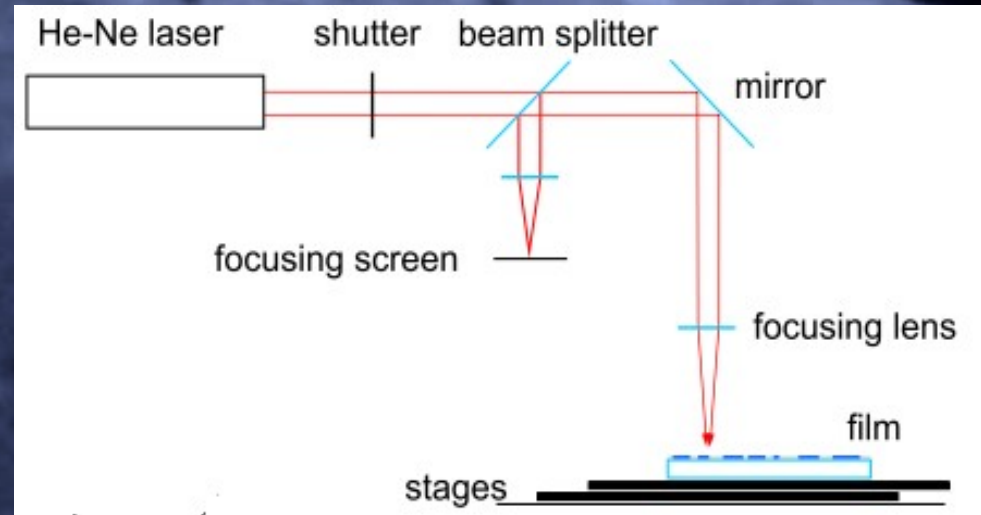
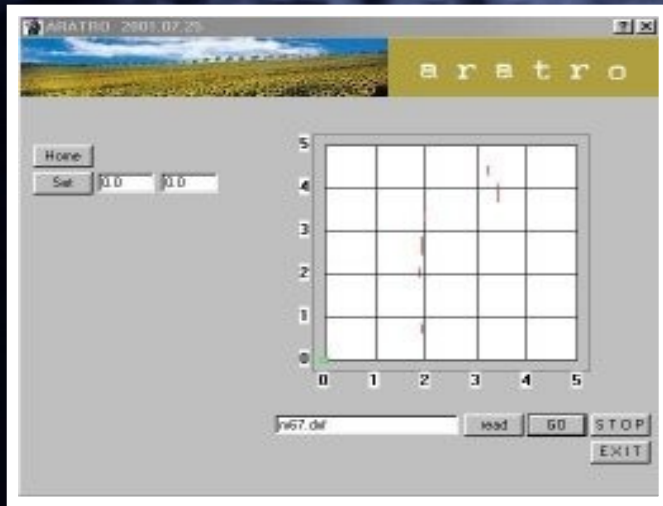
Rewritable Focal Plane Mask



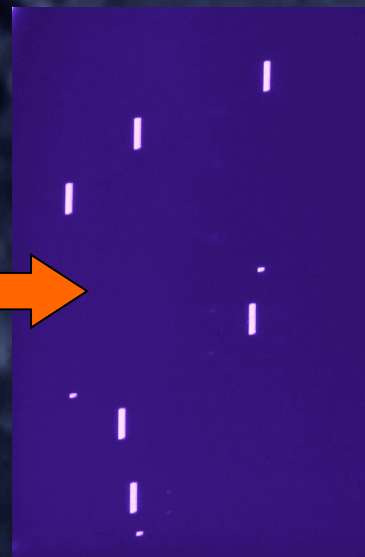
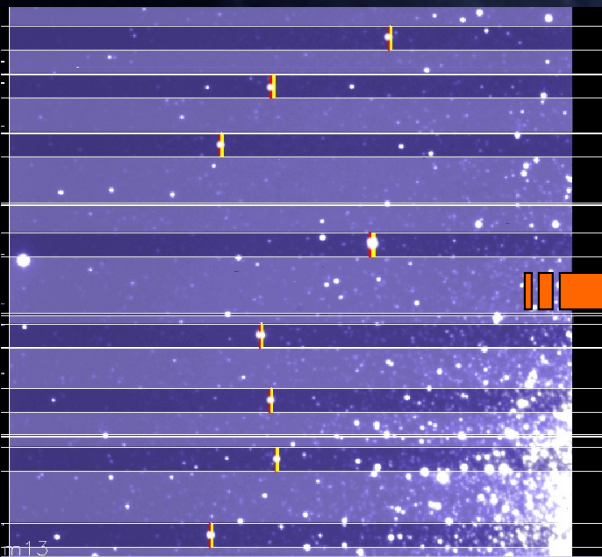
The colored form mimicks the metal and the uncolored form mimicks the slits.

Rewritable Focal Plane Mask

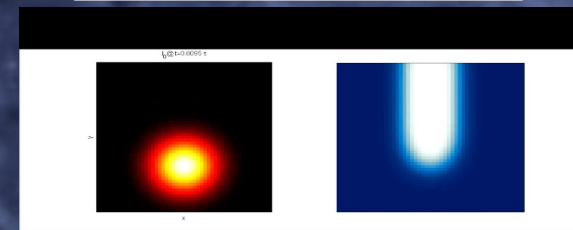
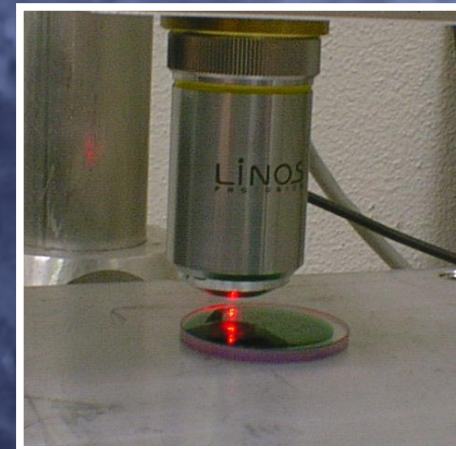
ARATRO (Plow): Writing software and hardware



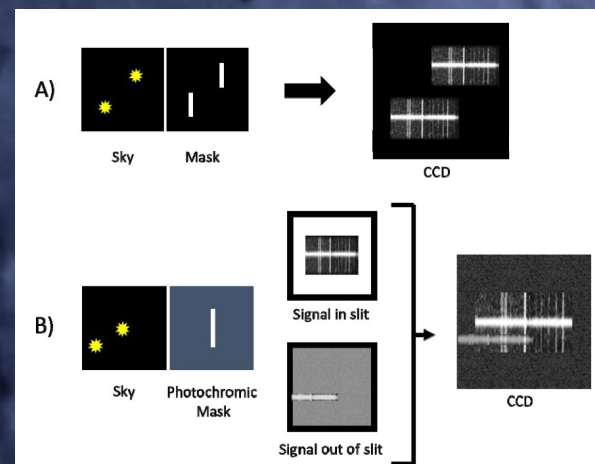
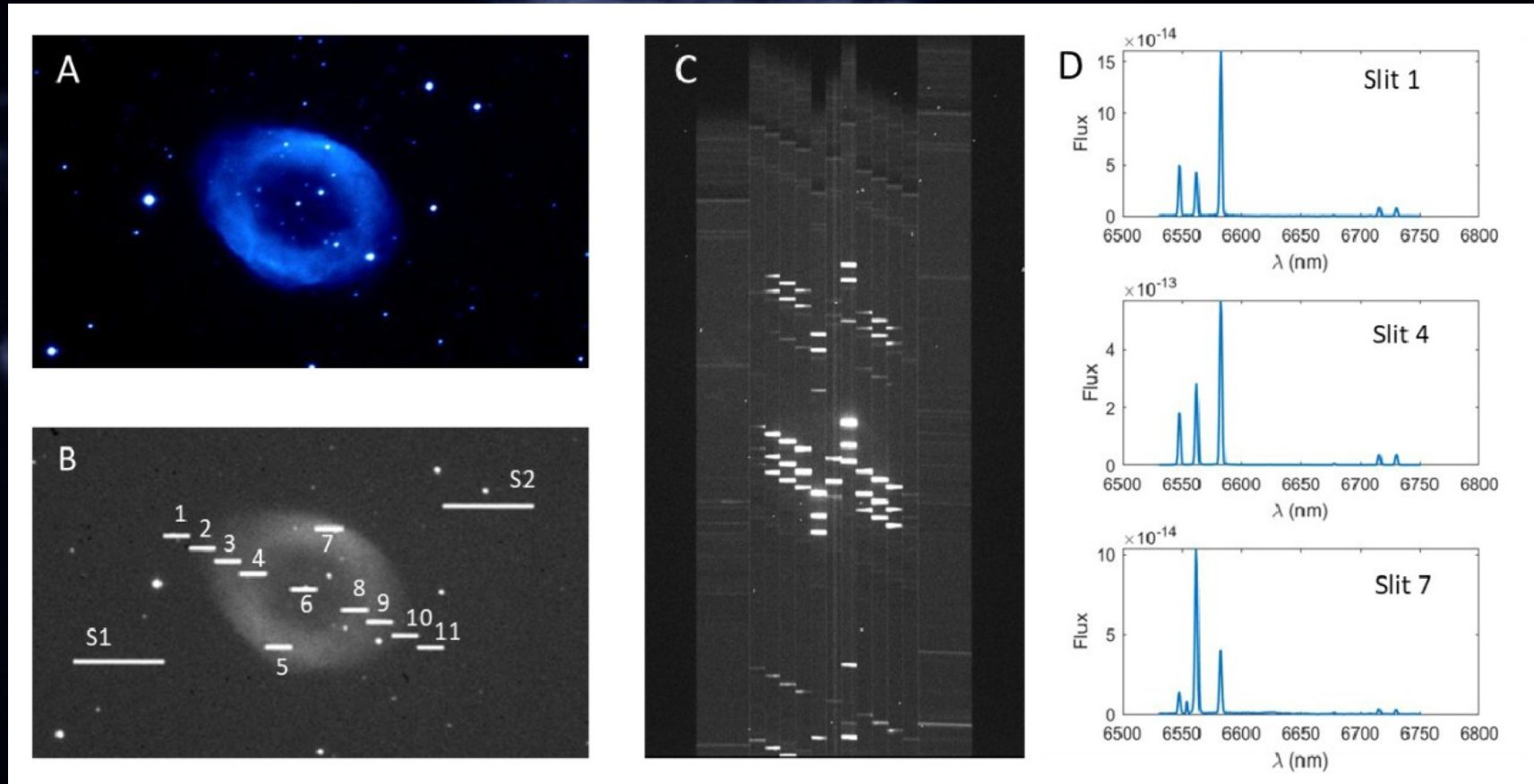
from the telescope...



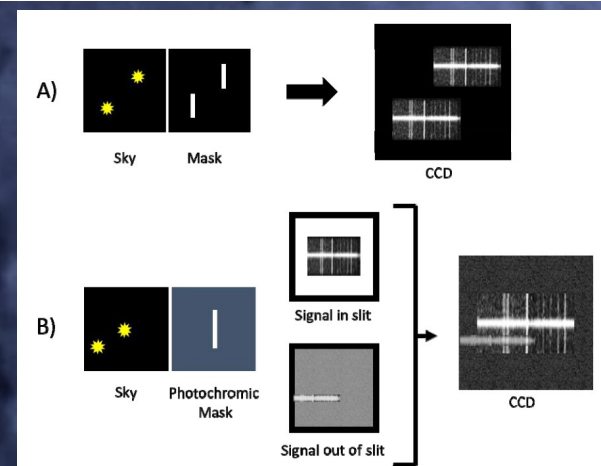
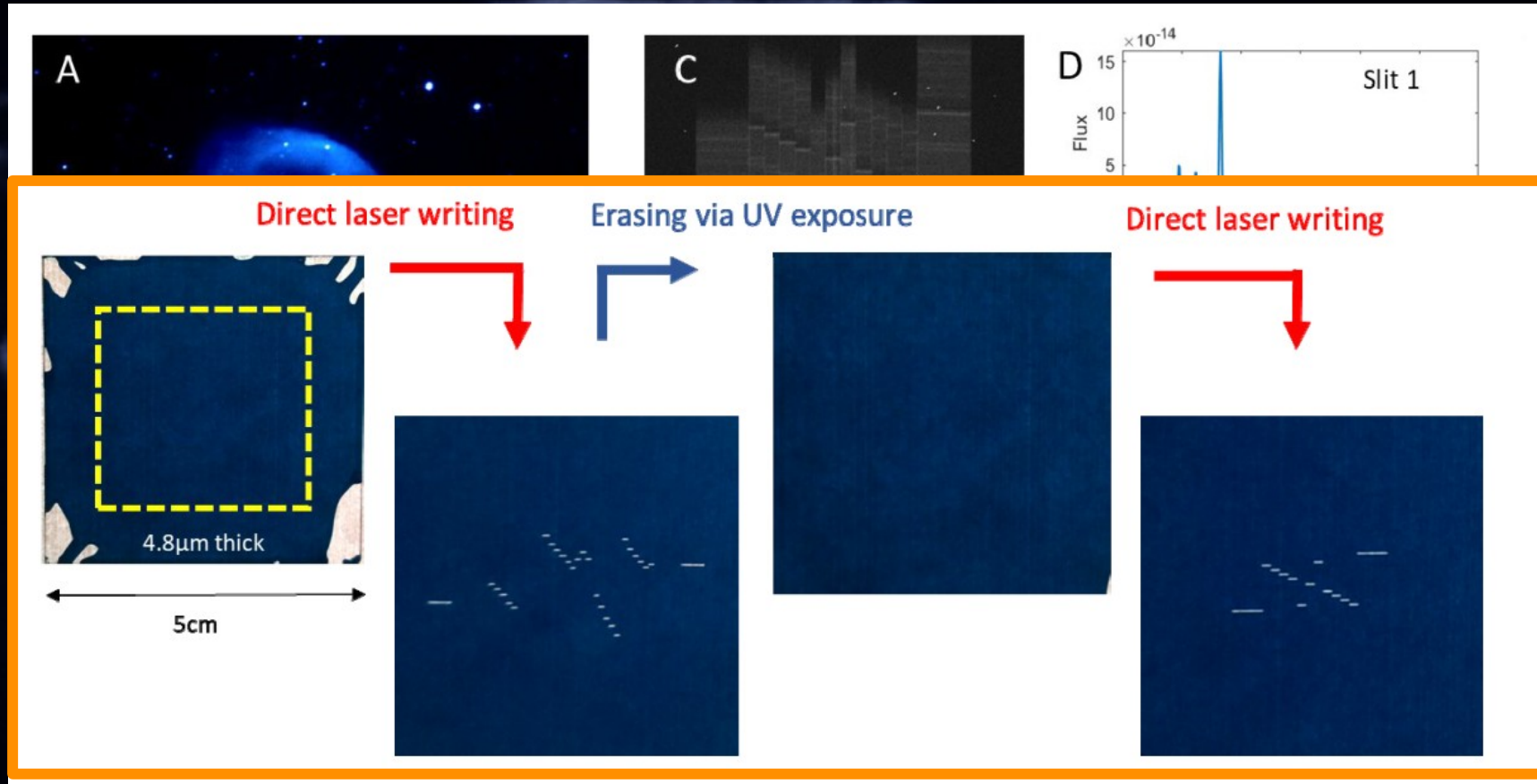
... to the mask



Rewritable Focal Plane Mask on sky



Rewritable Focal Plane Mask on sky

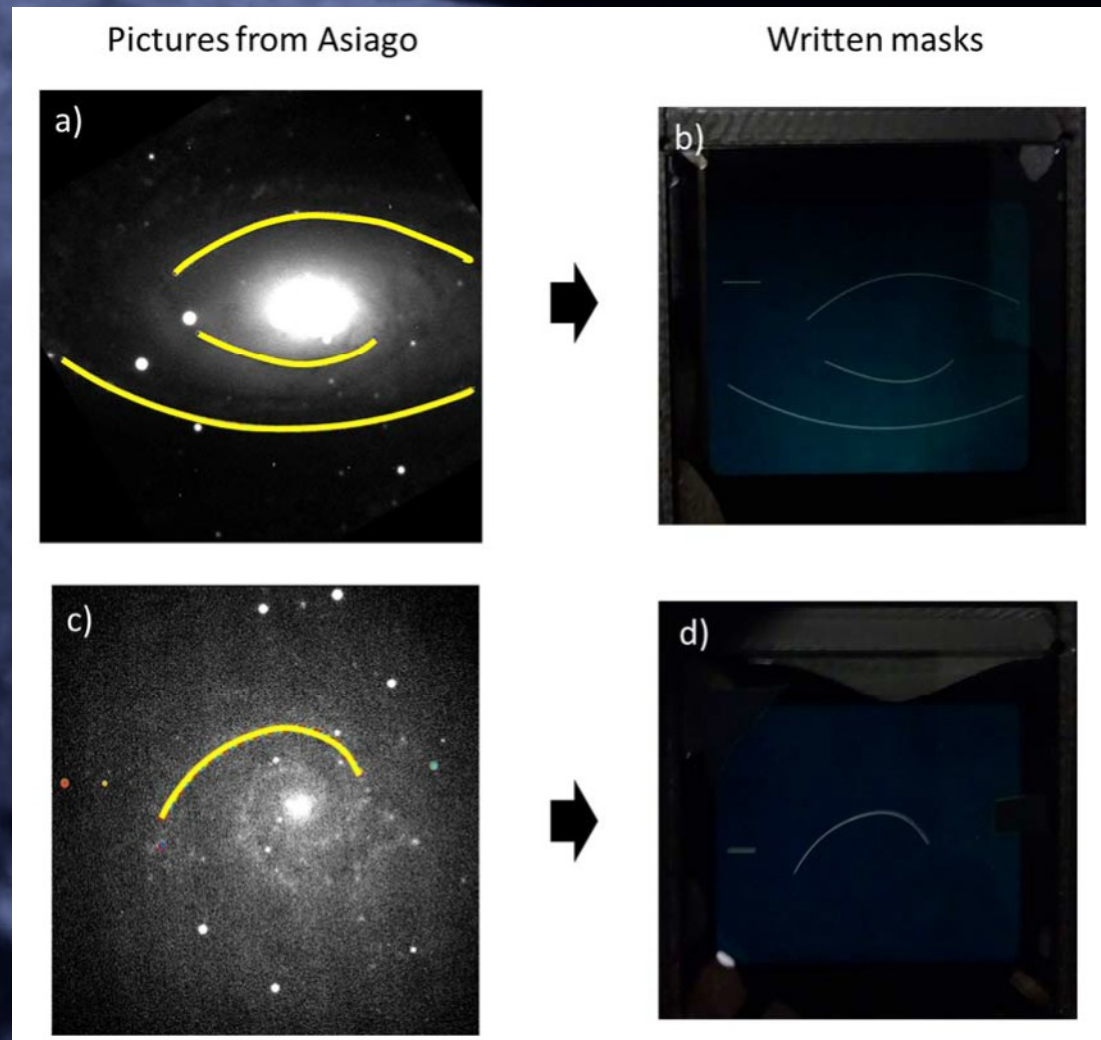


Rewritable Focal Plane Mask on sky



We can make the slit with the shape we want.

Curved slits "following" the galaxy arms.



Key Parameters

The key parameters for having an efficient device are:

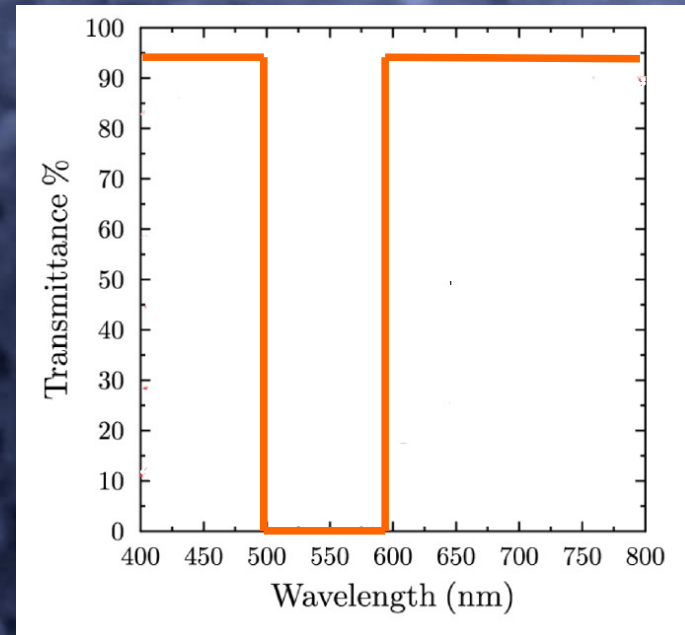
- Wavelength range;
- Contrast value;



Key Parameters: Wavelength range

The key parameters for having an efficient device are:

- Wavelength range;
- Contrast value;



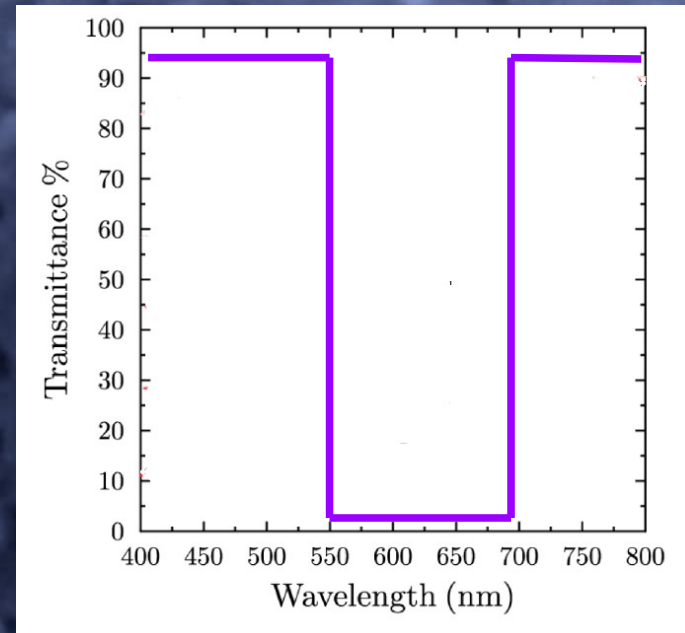
The wavelength range and position where the change in transparency occurs is important especially for the masks.

It means the possibility to collect spectra in a wider band.

Key Parameters: Wavelength range

The key parameters for having an efficient device are:

- Wavelength range;
- Contrast value;



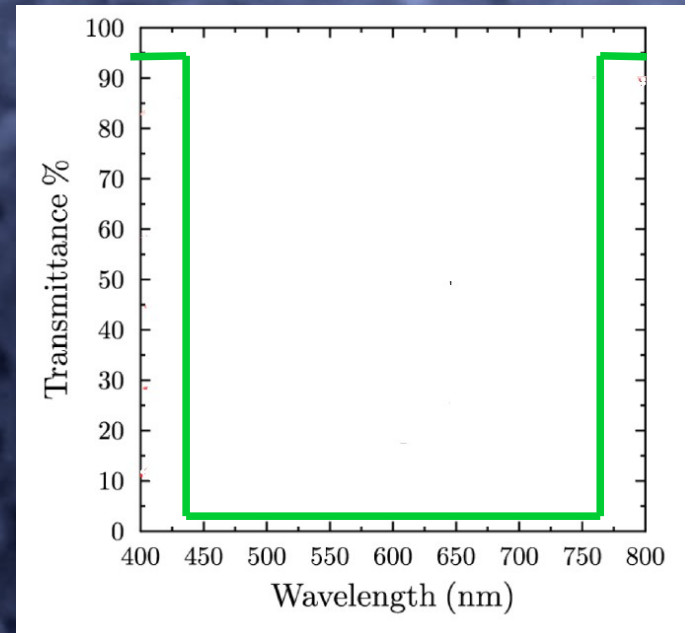
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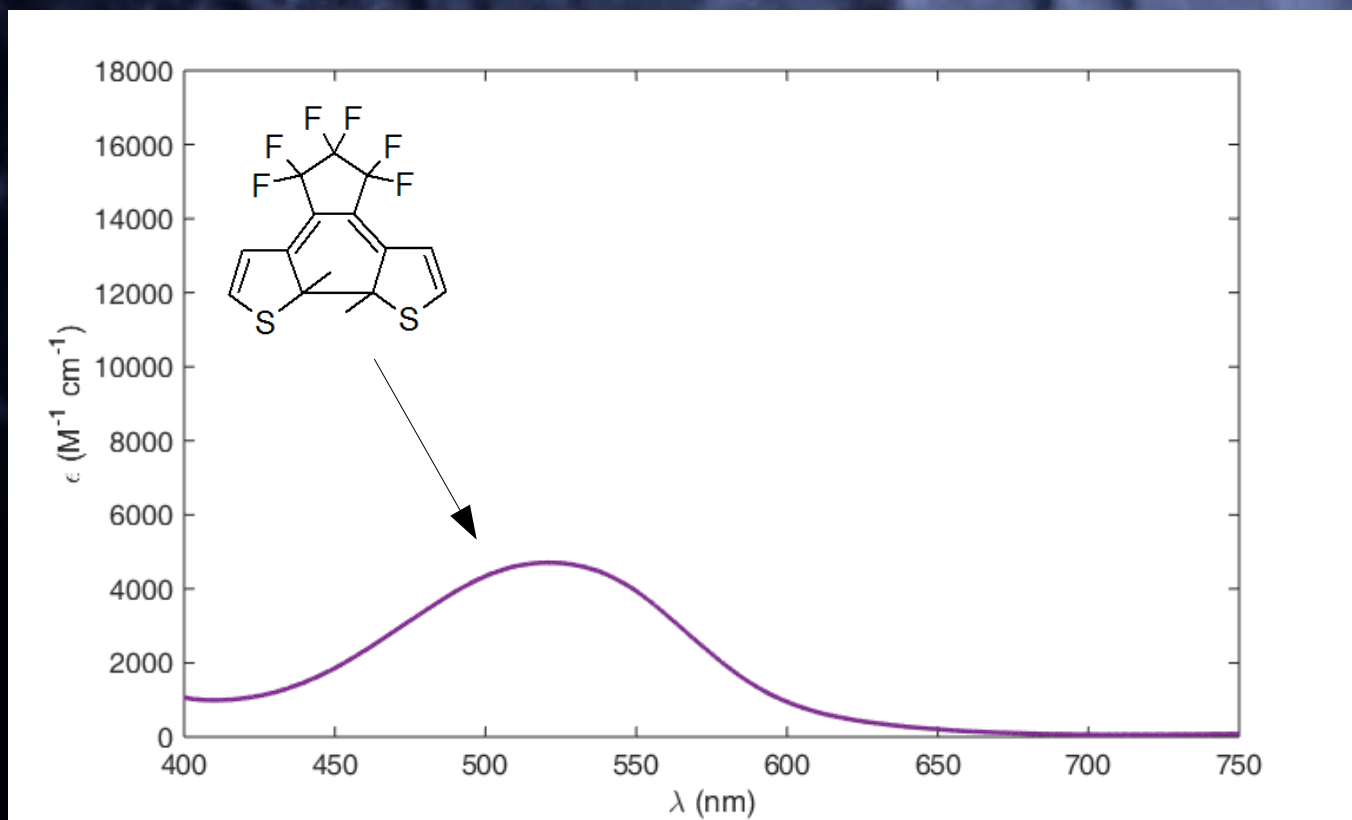


The wavelength range and position where the change in transparency occurs is important especially for the masks. It means the possibility to collect spectra in a wider band.

Key Parameters: Wavelength range

We can play with the chemical structure of the diarylethene:

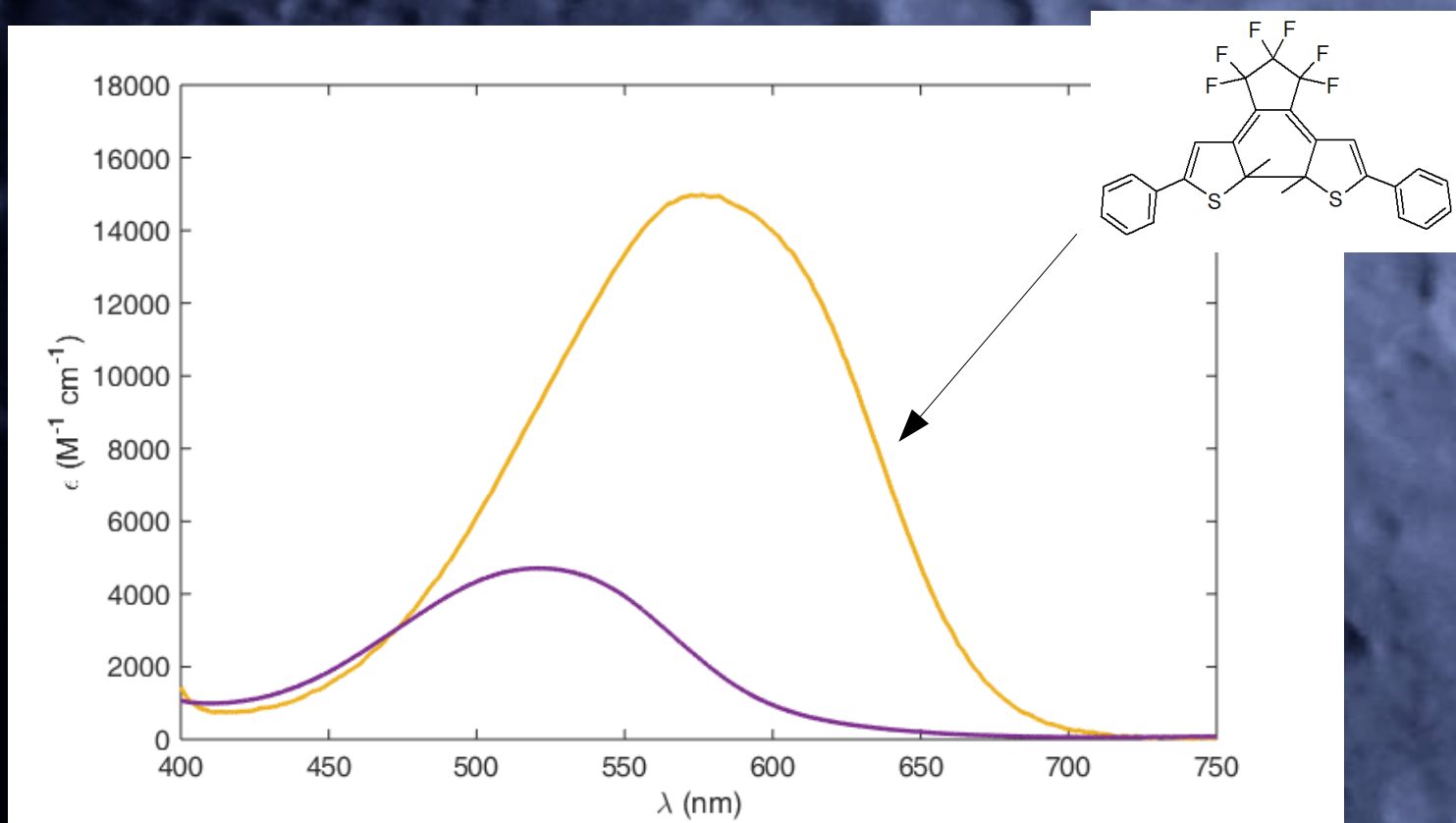
- Extension of π conjugation;
- Electroactive substituents (electron donor and acceptor)



Key Parameters: Wavelength range

We can play with the chemical structure of the diarylethene:

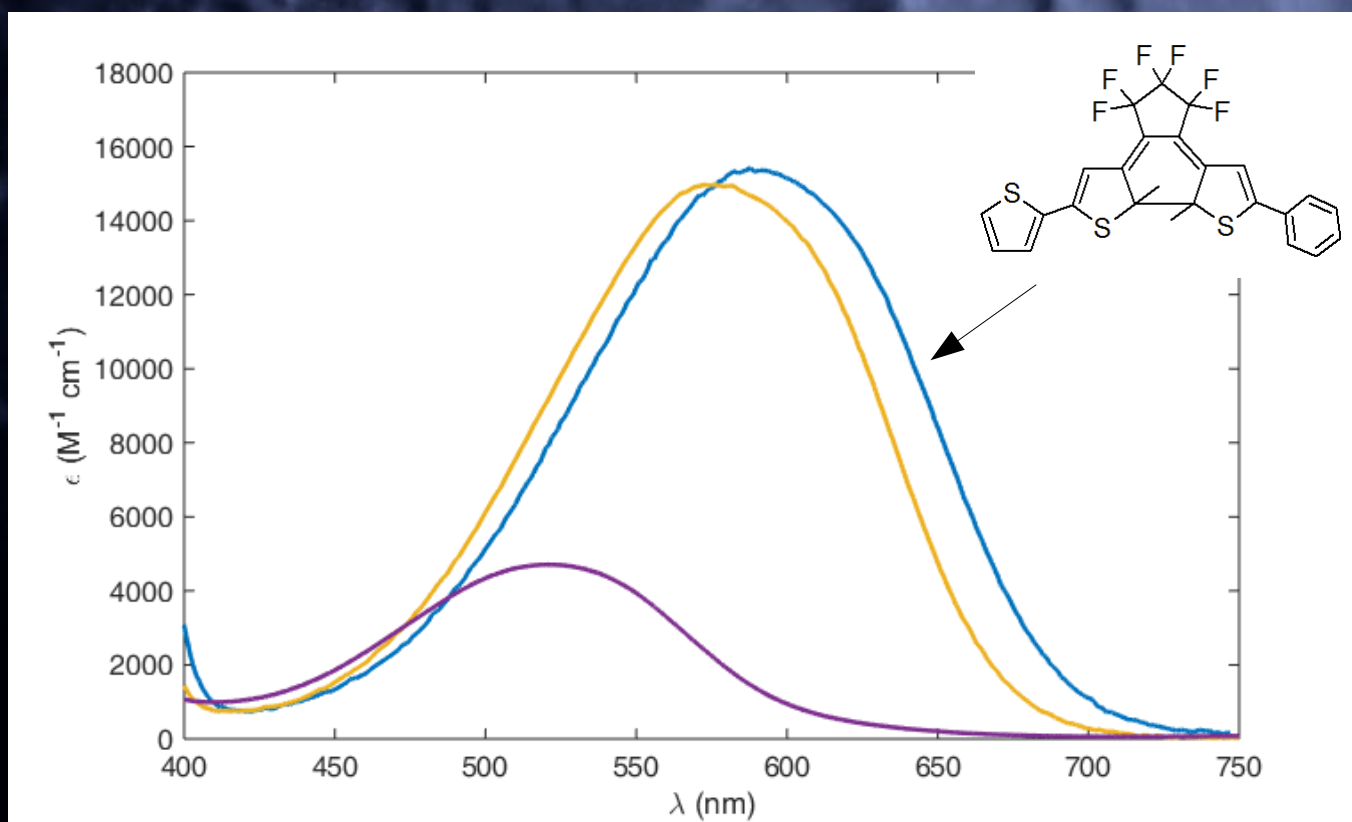
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Key Parameters: Wavelength range

We can play with the chemical structure of the diarylethene:

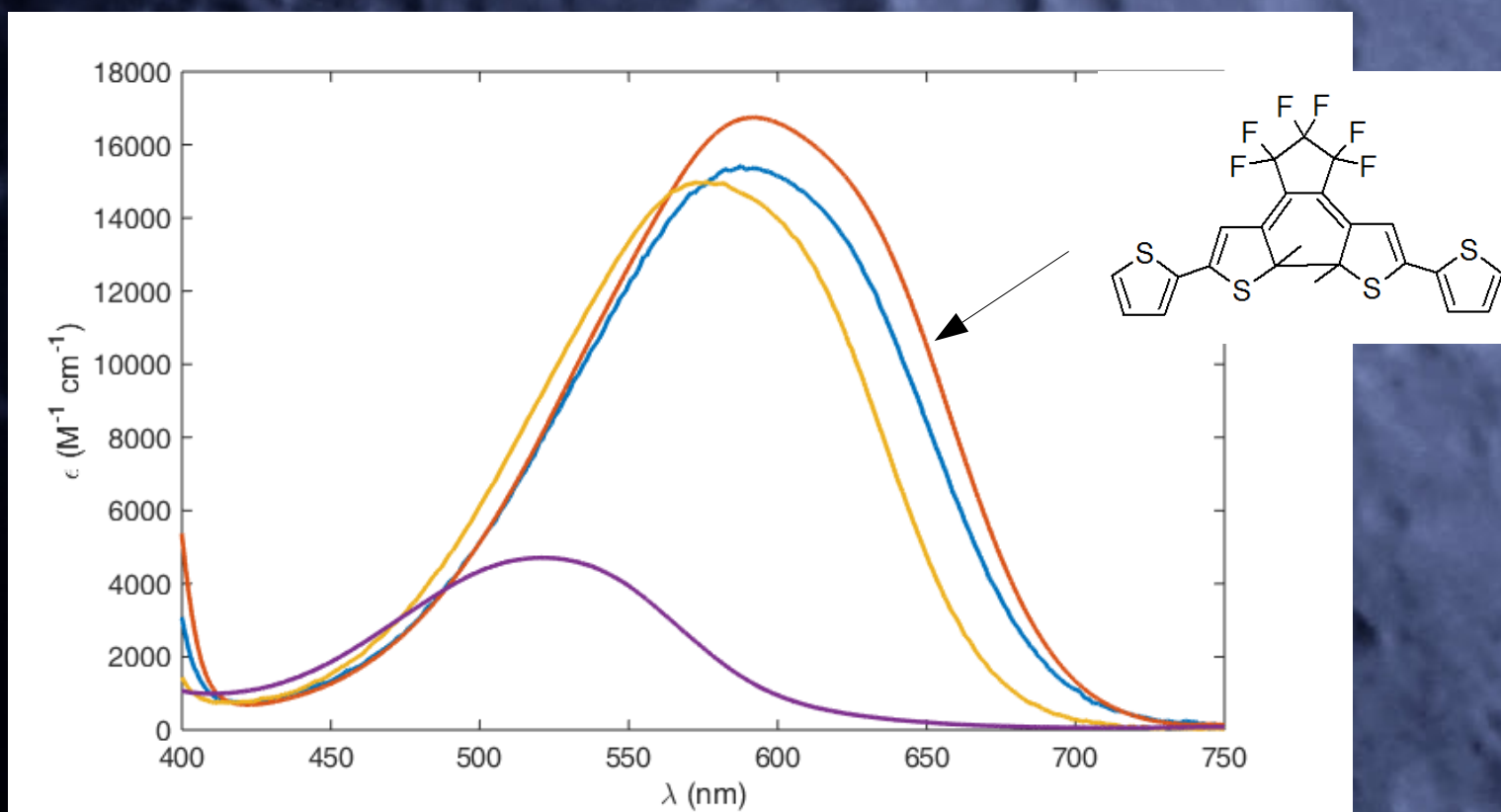
- Extension of π conjugation;
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Key Parameters: Wavelength range

We can play with the chemical structure of the diarylethene:

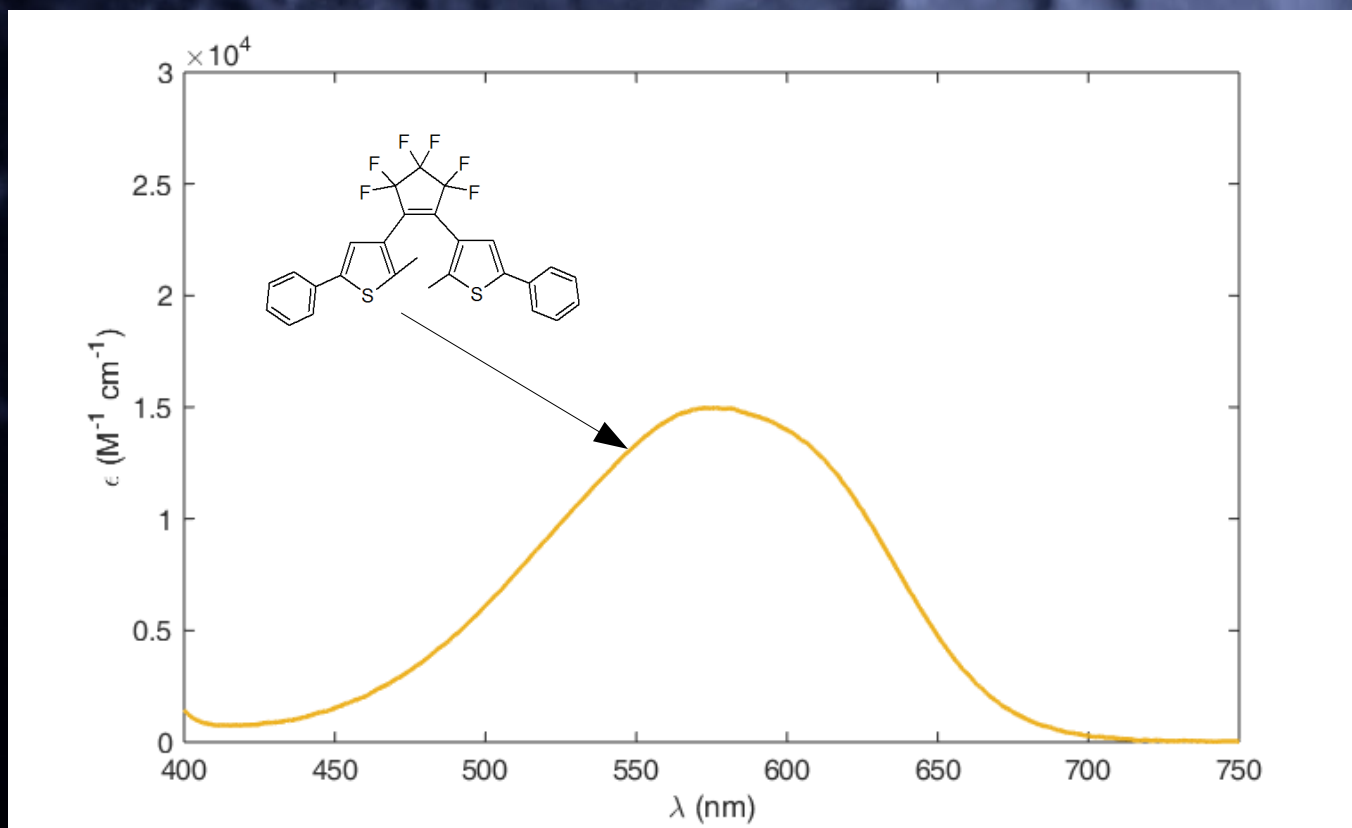
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Key Parameters: Wavelength range

We can play with the chemical structure of the diarylethene:

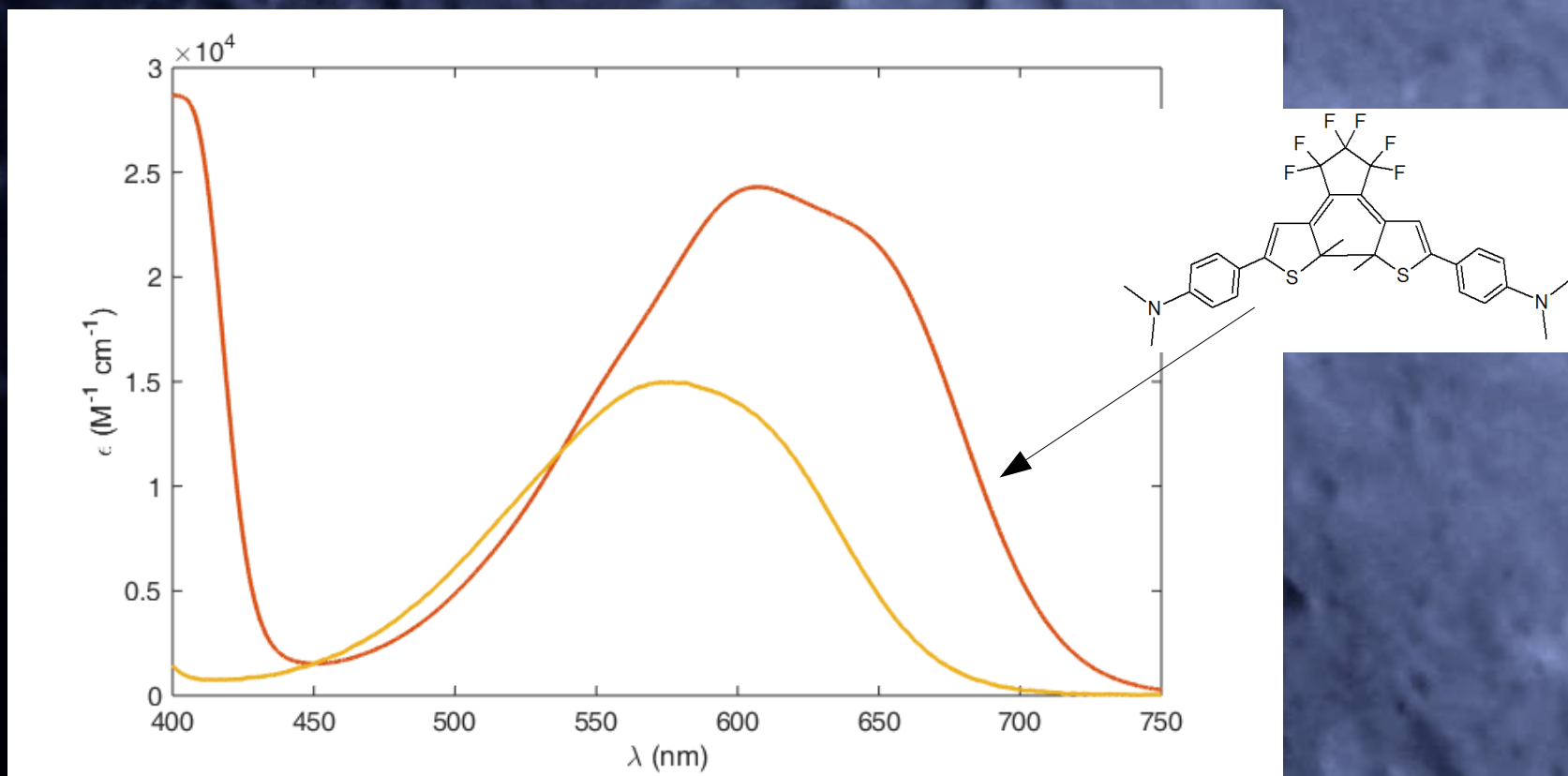
- Extension of π conjugation;
- Electroactive substituents (electron donor and acceptor)



Key Parameters: Wavelength range

We can play with the chemical structure of the diarylethene:

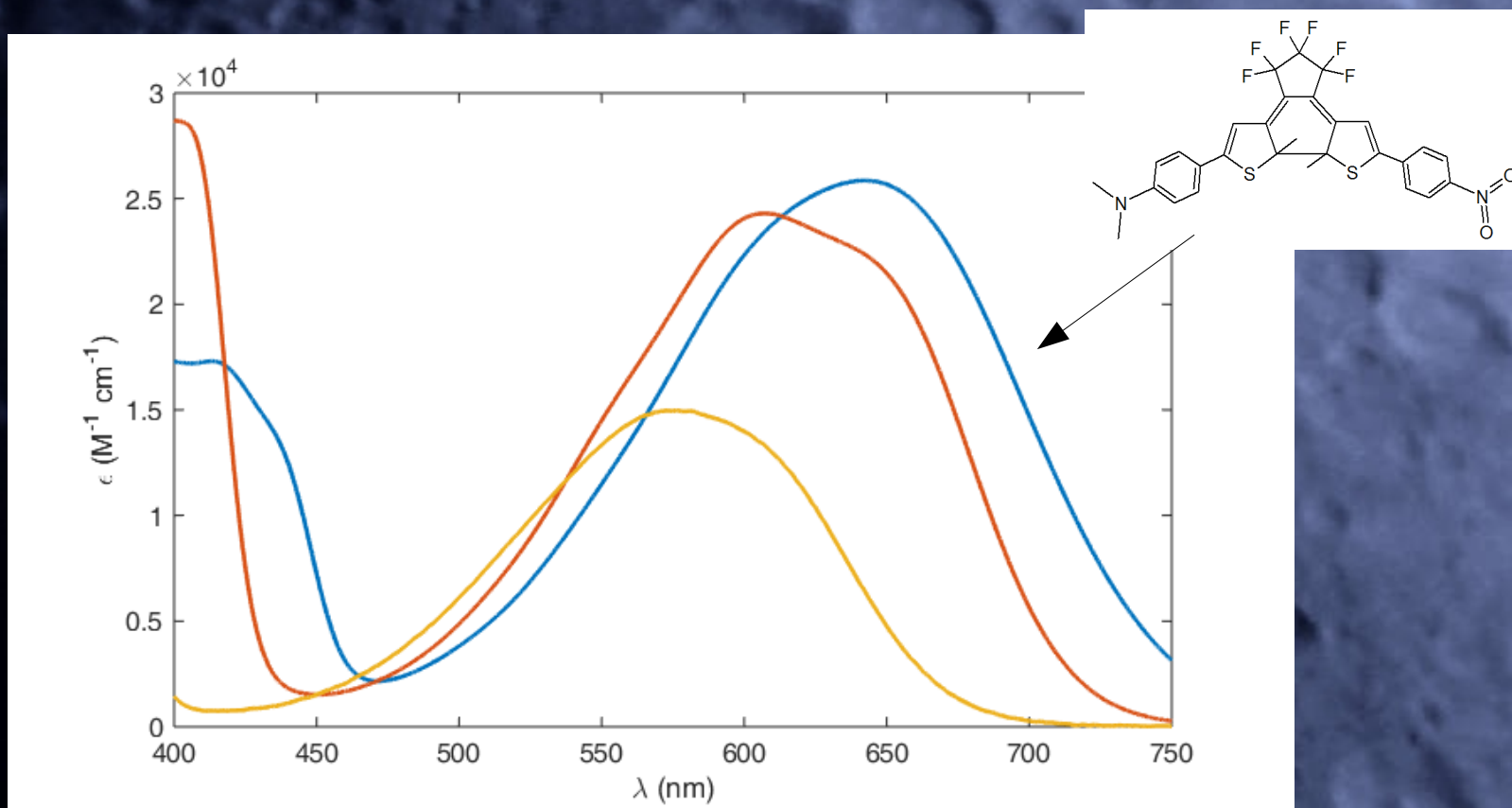
- Extension of π conjugation;
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Key Parameters: Wavelength range

We can play with the chemical structure of the diarylethene:

- Extension of π conjugation;
- Electroactive substituents (electron donor and acceptor)



Key Parameters: Wavelength range

With the help of good chemists, we can cover the whole visible spectrum combining different diarylethenes

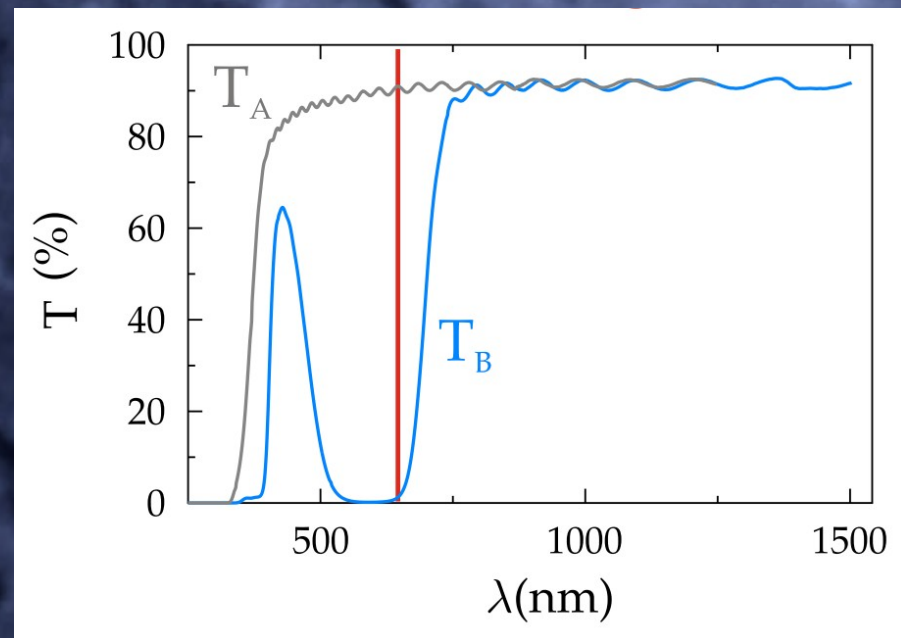


Key Parameters: Contrast

The key parameters for having an efficient device are:

- Wavelength range;
- Contrast value;

$$CT(\lambda) = \frac{T_A}{T_B}$$



The contrast CT is defined as the ratio of the transmittance in the uncolored form and the colored one

Key Parameters: Contrast

The key parameters for having an efficient device are:

- Wavelength range;
- Contrast value;



$$CT(\lambda) = \frac{T_A}{T_B}$$

In the case of focal plane masks, the higher the **CT**, the lower the sky contamination (S/N). The high contrast is required over a wide wavelength range.

Key Parameters: Contrast

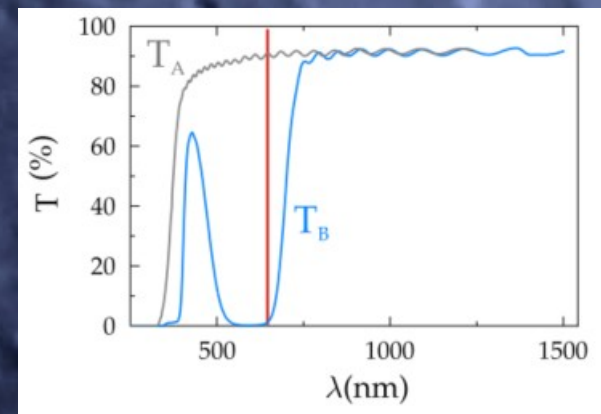
The key parameters for having an efficient device are:

- Wavelength range;
- Contrast value;

$$CT(\lambda) = \frac{T_A}{T_B}$$

$$T_B = 10^{-\epsilon_B CD}$$

Ext. Coeff. Concentration Thickness



The contrast is driven by the transmittance of the colored T_B form through the Lambert-Beer law.

Moving from molecule to material...

The key parameters for having an efficient device are:

- Wavelength range;
- Contrast value;

$$T_B = 10^{-\epsilon_B CD}$$

Ext. Coeff. Concentration Thickness



Make thick film with a large content of photochromic molecule!

Easy!

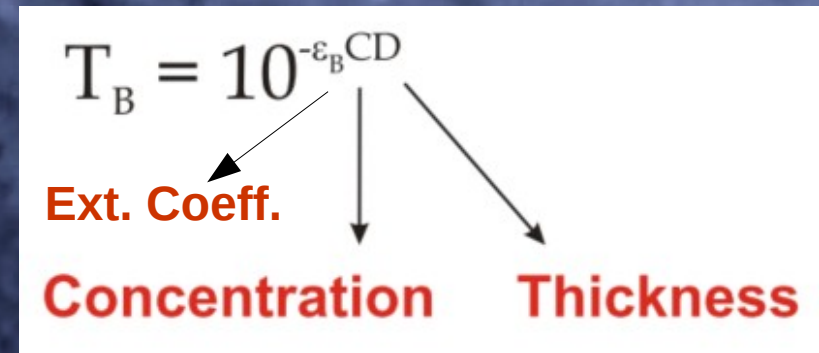
Moving from molecule to material...

The key parameters for having an efficient device are:

- Wavelength range;
- Contrast value;

$$T_B = 10^{-\epsilon_B CD}$$

Ext. Coeff. Concentration Thickness

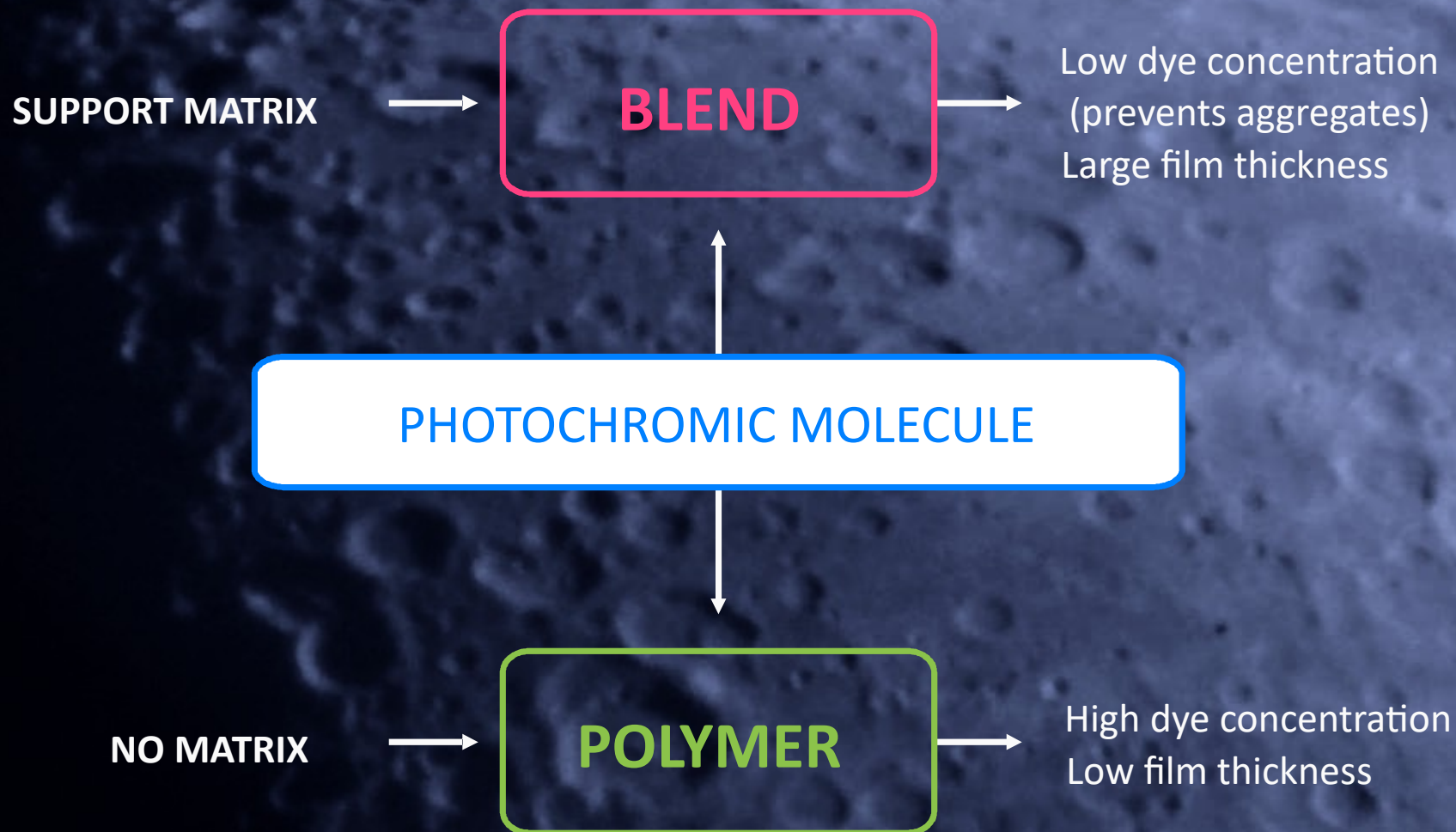


Yes, but how?

We have just the molecules
at the moment...



Strategies



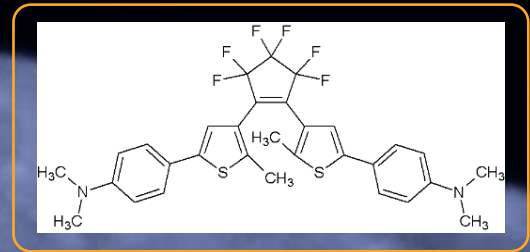
Blend

Very simple film production

Limited dye concentration to prevent segregation

Strong property modulation is required in the diarylethene moiety

Check the solubility of the molecule in the binder!



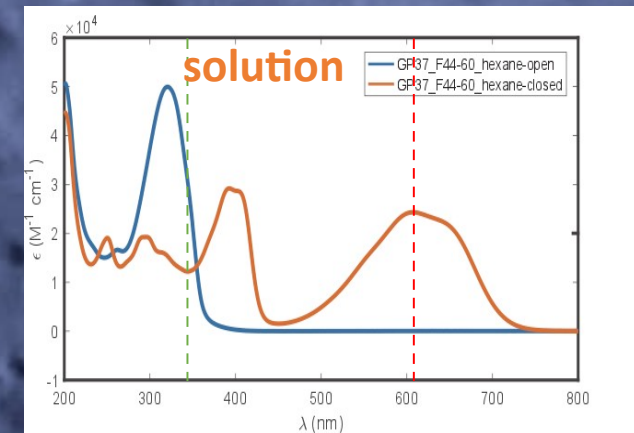
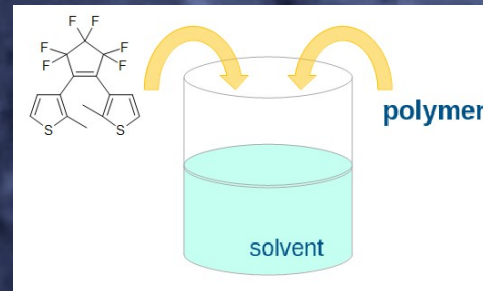
Polymer matrices (binders):

Polystyrene

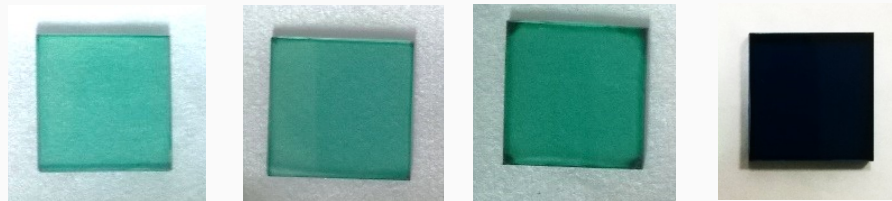
PMMA

Zeonex

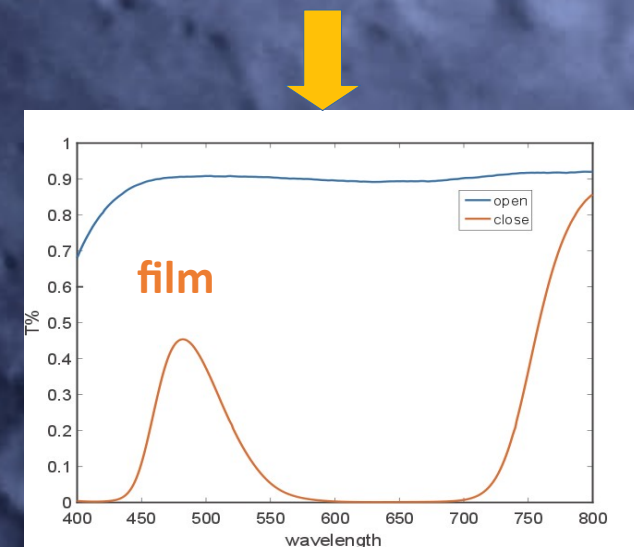
CAB (Cellulose Acetate Butyrate)



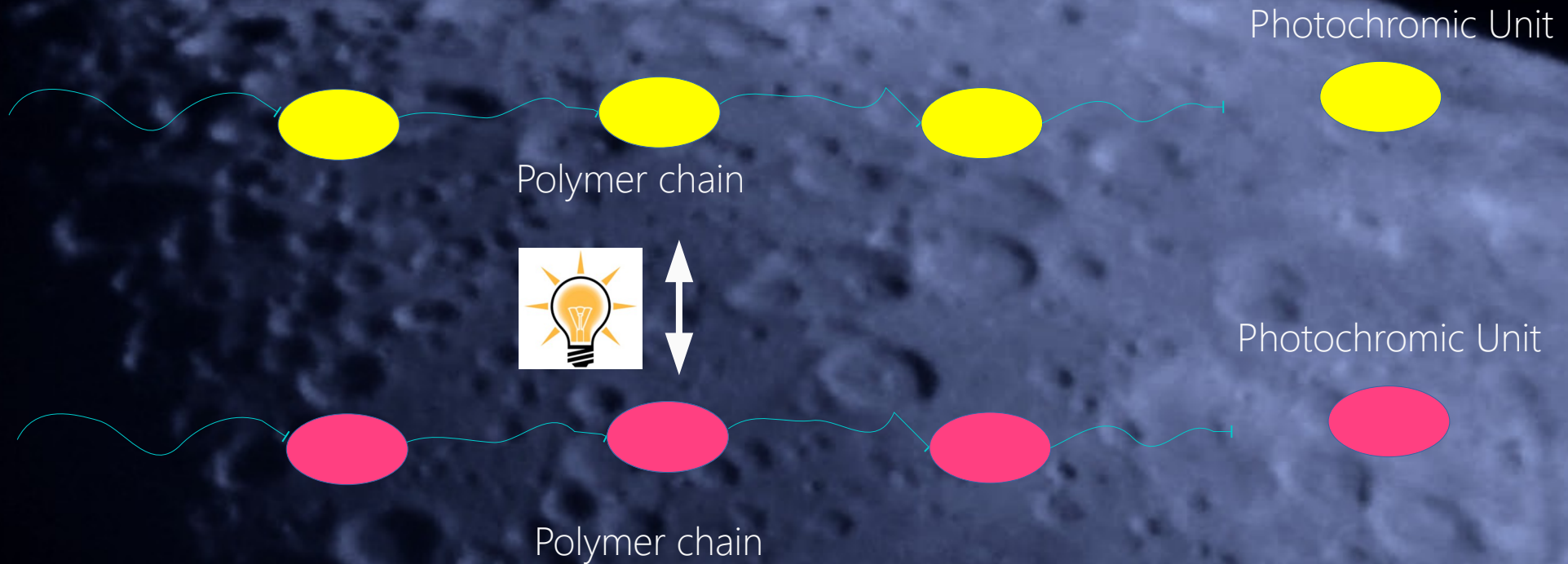
The molecule must be carefully chosen!



d (μm)	2.0	3.0	4.0	3.5
C (%wt.)	5	5	5	20
Abs@633nm	0.5	0.7	1.0	3.2

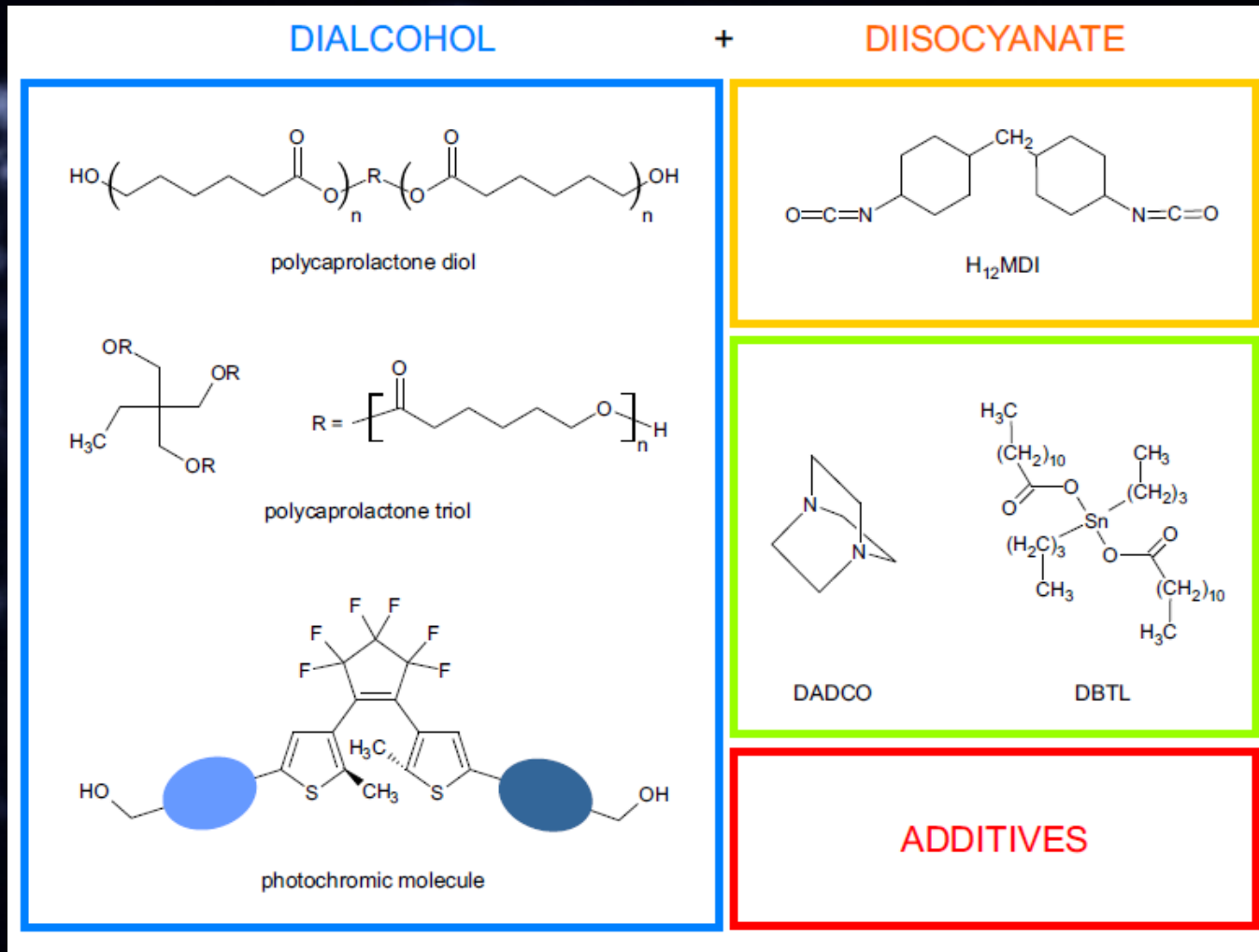


Backbone photochromic polymers

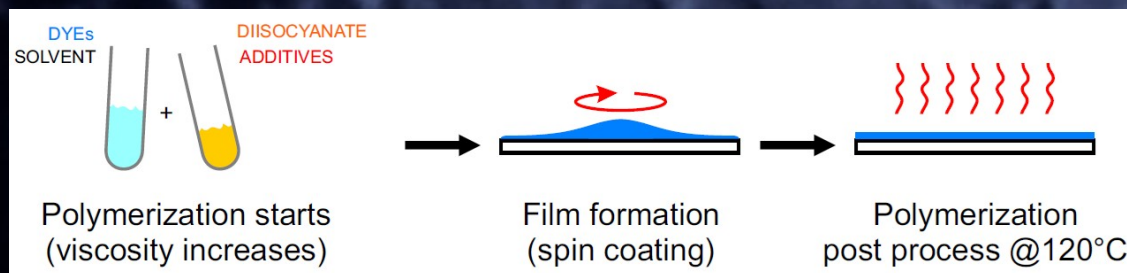


The monomer is photochromic, therefore it need reactive groups;
The optimization of the chemical structure is more difficult;
It is **easy to have high content photochromic films** without segregation.

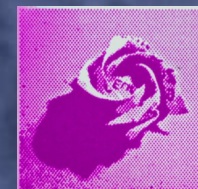
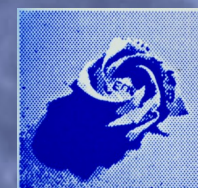
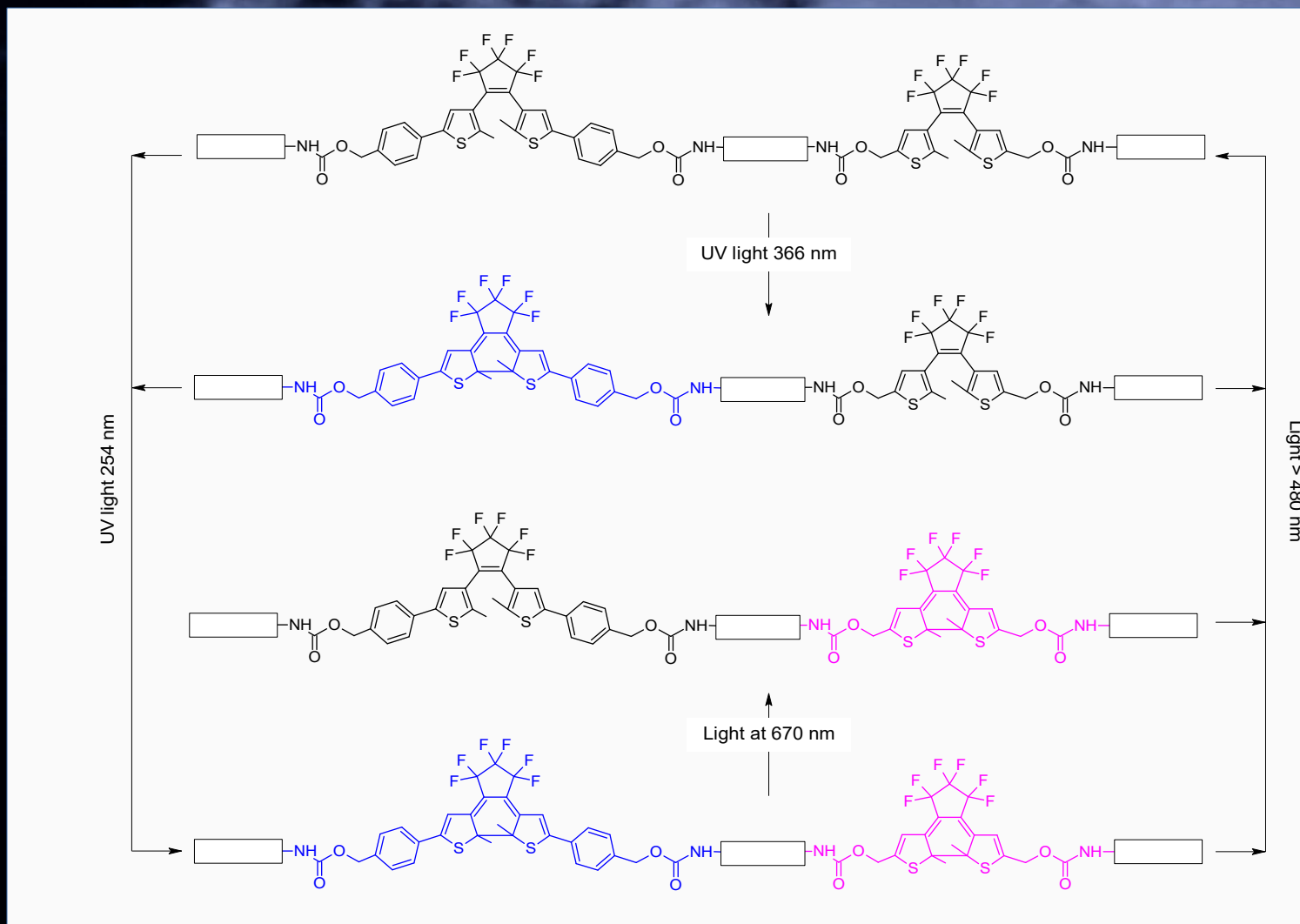
Photochromic polyurethane



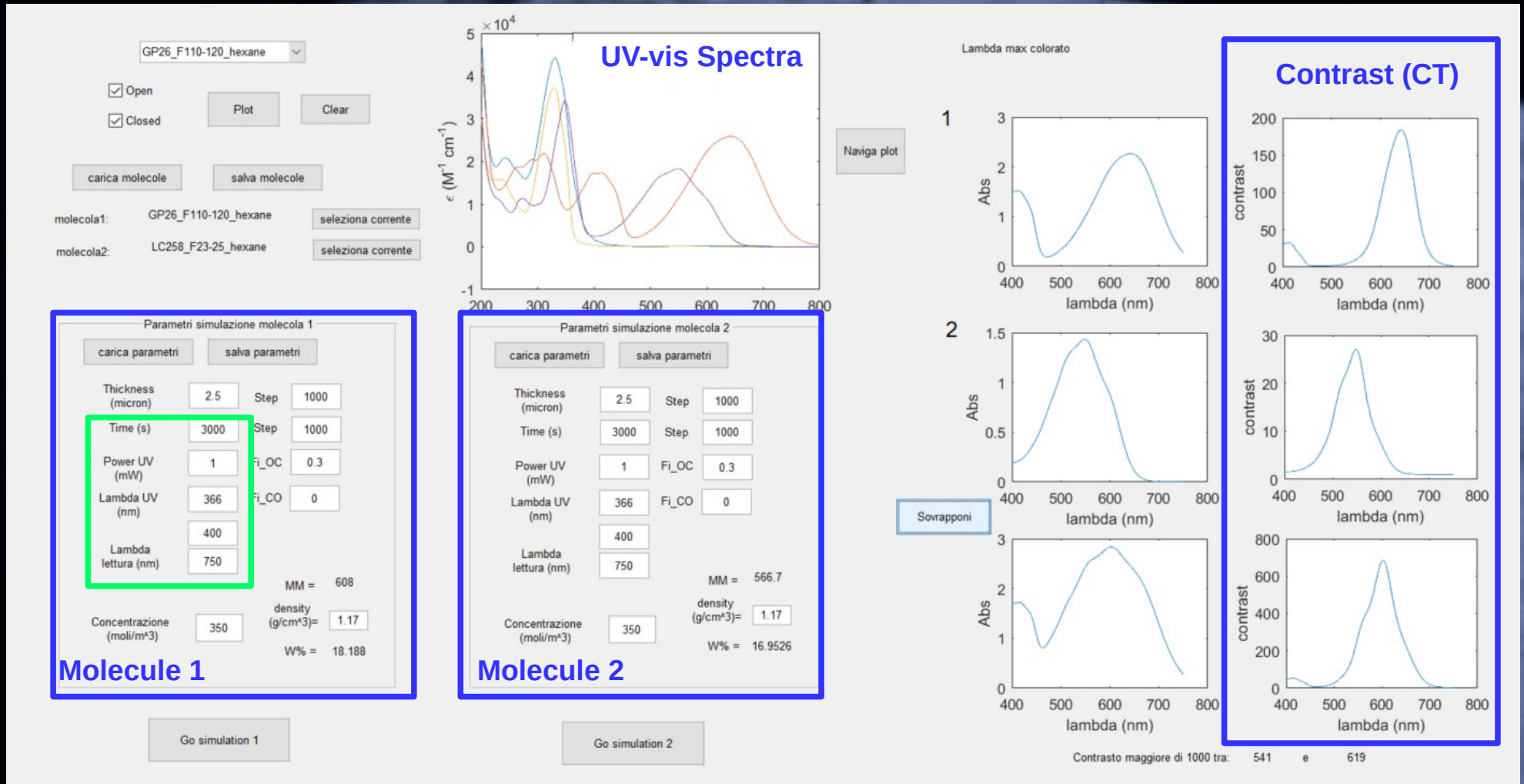
Processing (*in situ* polymerization)



Photochromic polyurethane

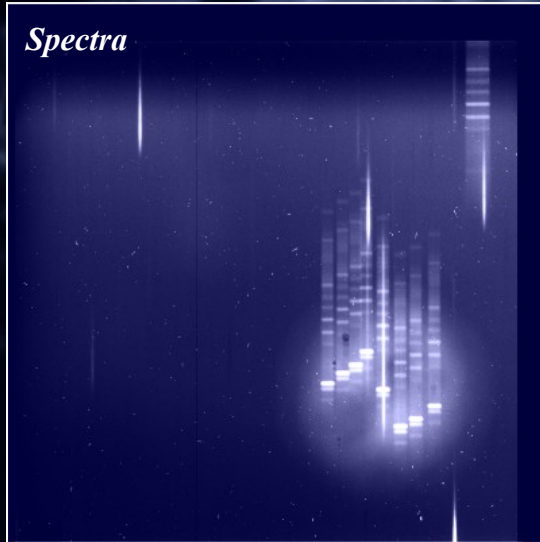


A designing tools for photochromic films



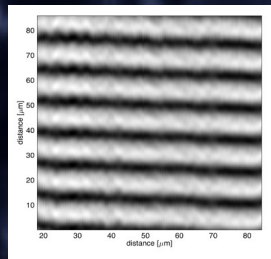
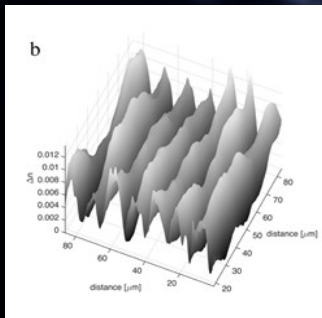
We developed a tool for the design of films with a target value of contrast as function of the "chemistry".

What we did with photochromic materials



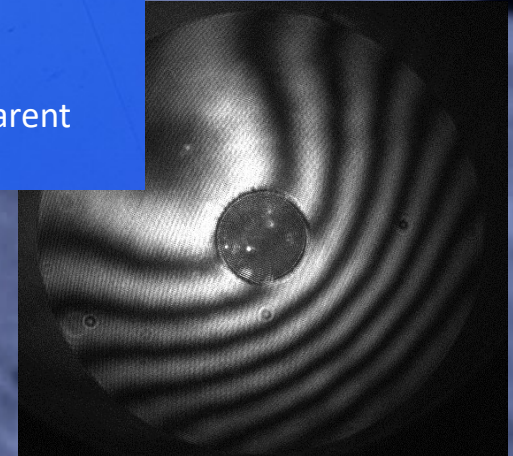
Focal Plane Masks

Bianco A., *Astron. Nachr.*, **326** (5) pp. 370–374, 2005
Luca Oggioni, et al., *Proc. SPIE* 10706, 1070636 (2018)



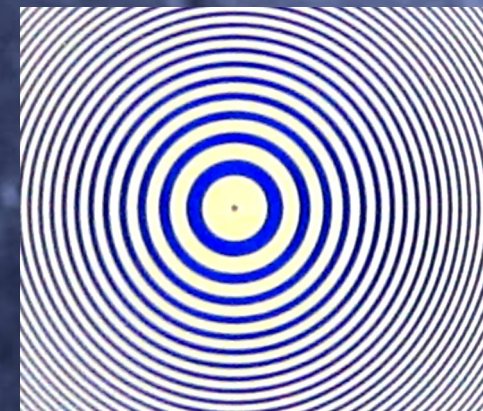
Mask for optical lithography

Pariani, G.; et al *Opt. Lett.* **38**, 3024 (2013).
G. Pariani, et al. *Adv. Mater. Tech.* **3**(3),1700325 (2018)



Point Diffraction Interferometer

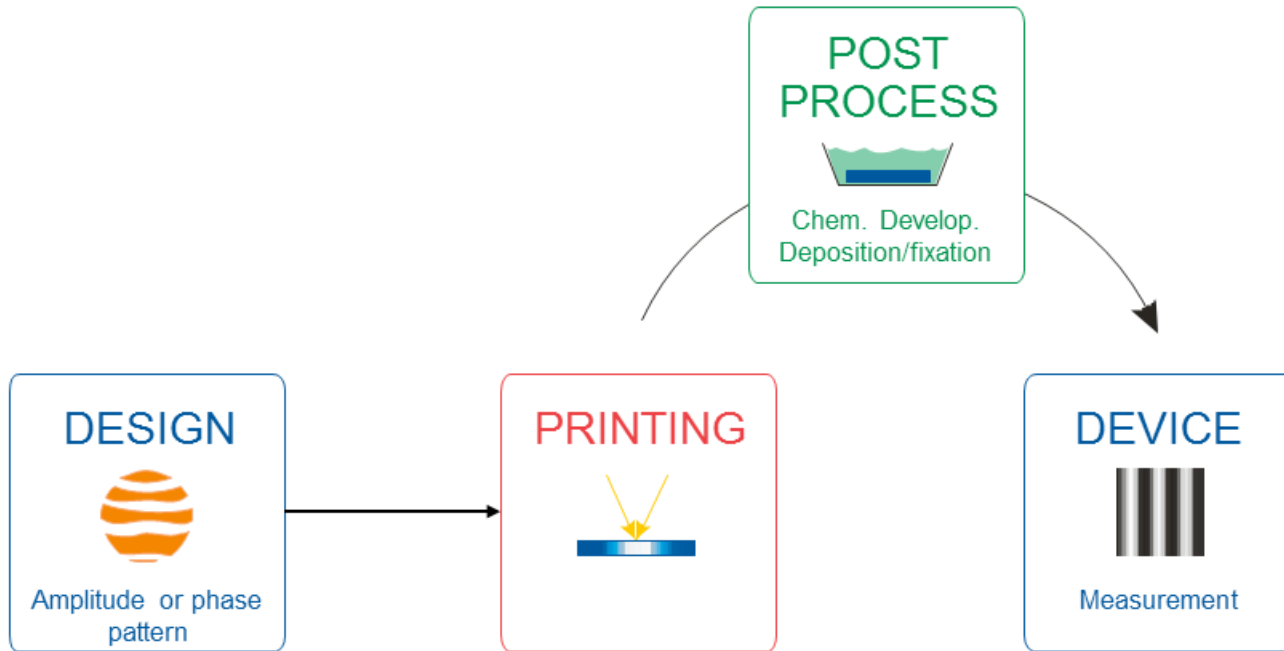
Quintavalla, M.; et al *Optics and Lasers in Engineering* **2014**, *56*, 134.



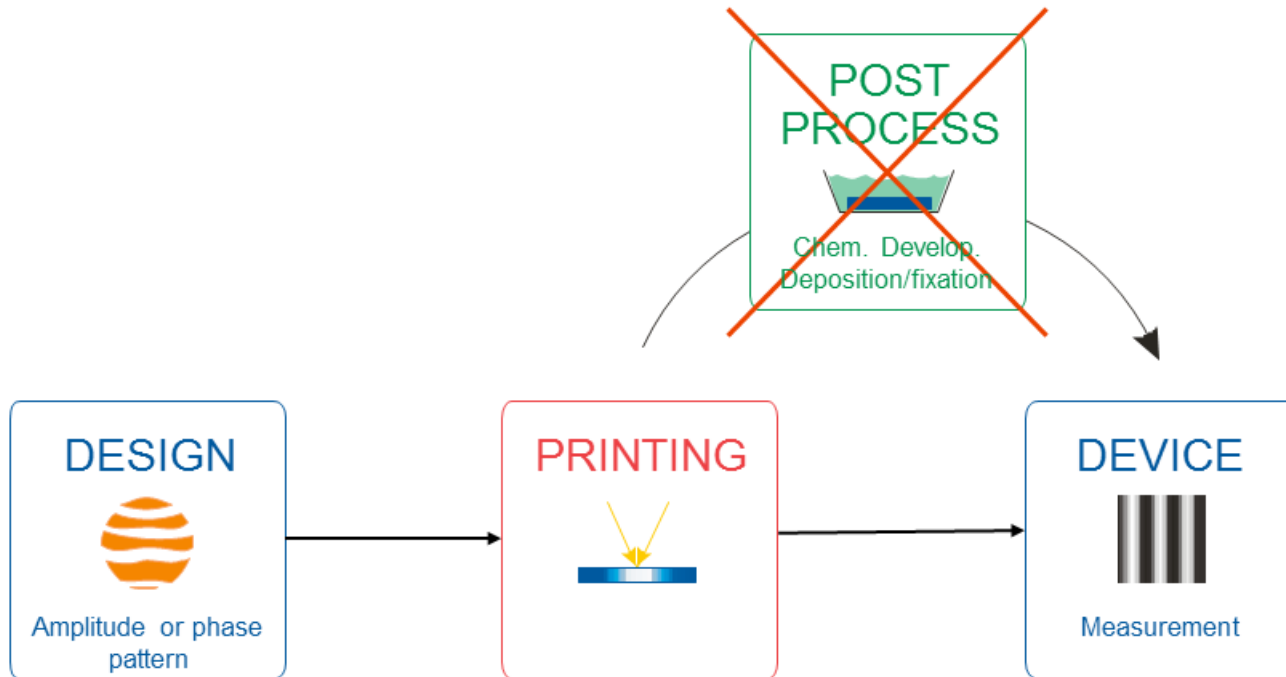
CGH for optical testing

Pariani, G.; et al *Opt. Express* **19**, 4536 (2011)
Alata, R. et al, *Proc. SPIE.* **9912**, 991234 (2016)
F. Zamkotsian, et al. *Opt. Express* **27**(19), 26446 (2019).
I. Oggioni et al *Materials* **12**(7) 2810 (2019)

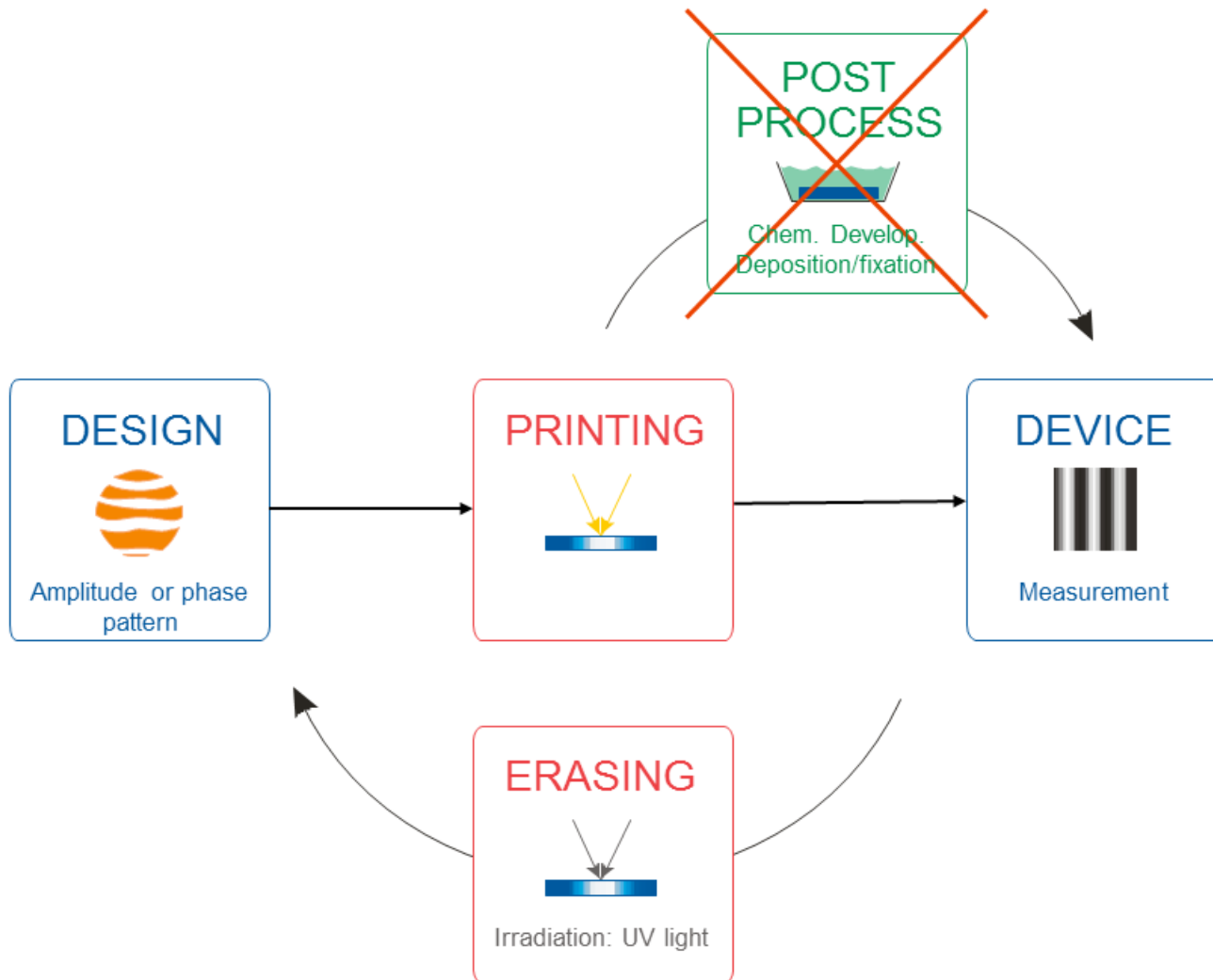
Simplify the process...



Simplify the process...



Simplify the process...



We developed a simple and reconfigurable platform for masks, waveguides, holograms,...

Dispersing elements

Spectral Resolution

Low Resolution (100 – 2000);
Medium Resolution (2000 – 5000);
High Resolution (5000 – 50000);
Very High Resolution (50000 – 200000).

Spatial Resolution

Single Object;
Multi Object Spectroscopy (MOS);
Spectral imaging.

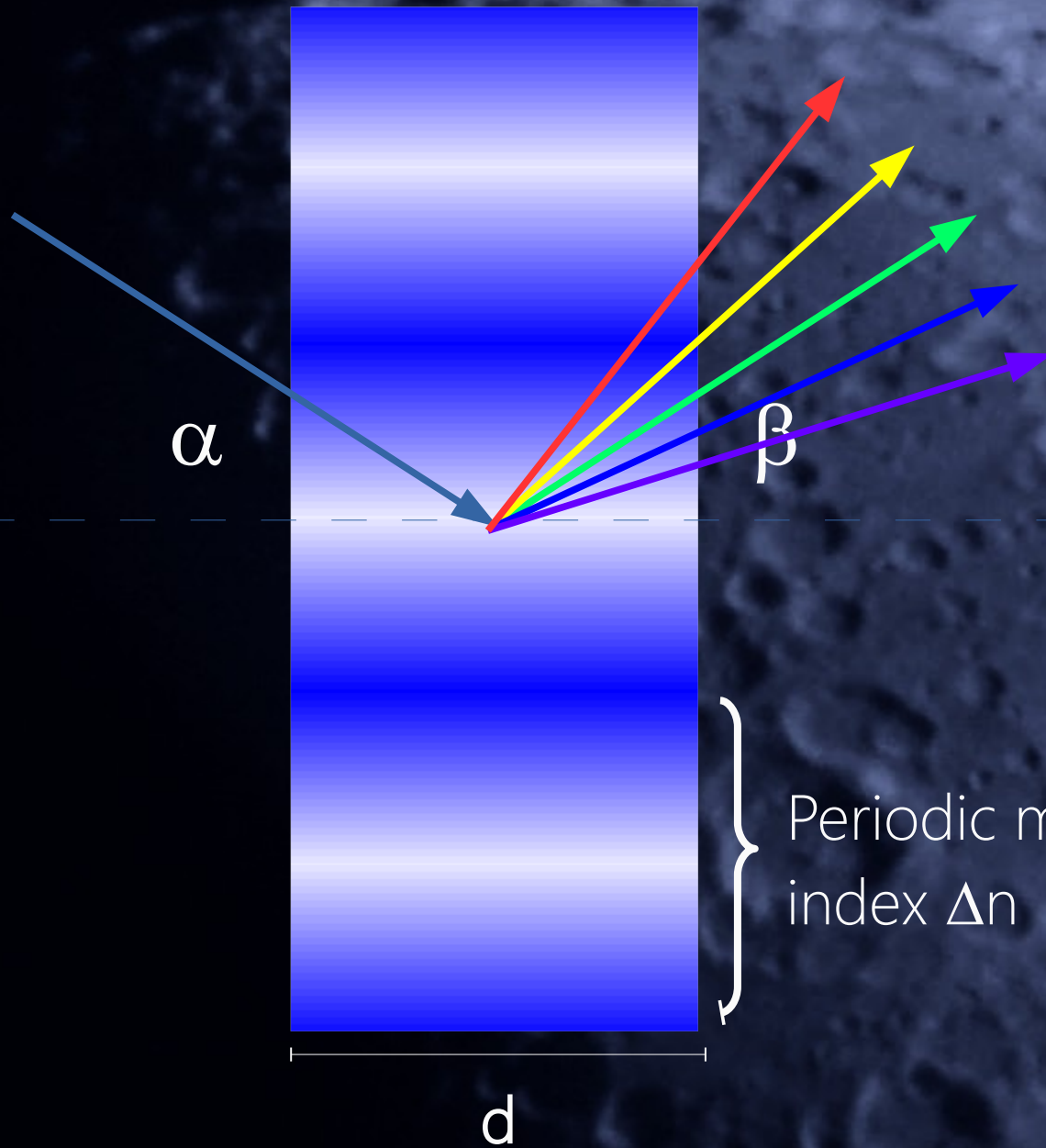
Spectral range

UV-Vis (0.3 – 1.0 μm)
NIR (0.9 – 2.5 μm)
MIR (2.5 – 25 μm)

**Dispersing element (grating)
is a key element**



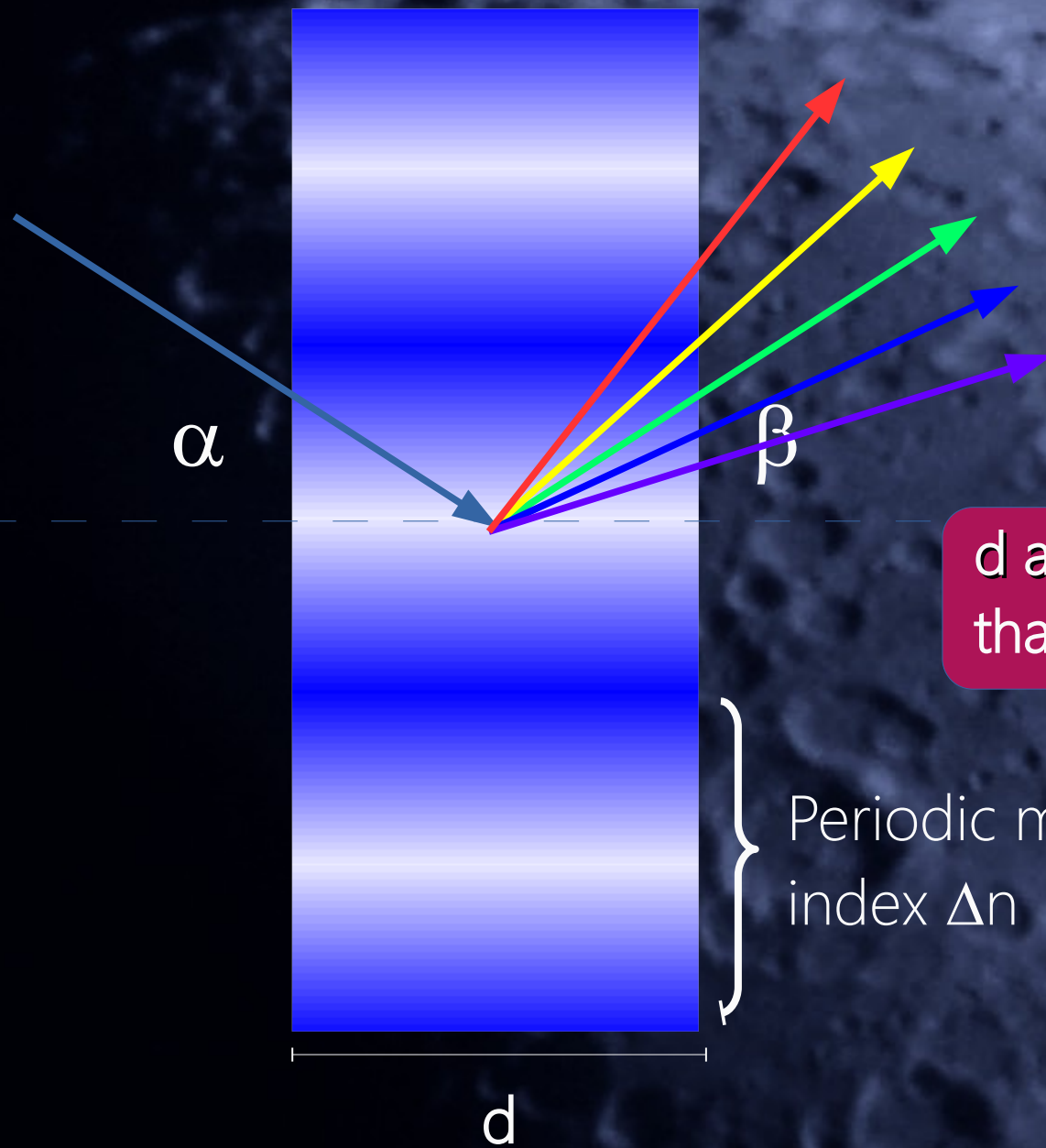
Dispersing elements: VPHG



The diffraction occurs thanks to a periodic modulation of the refractive index in the volume of the material.

Periodic modulation of the refractive index Δn (usually sinusoidal).

Dispersing elements: VPHG



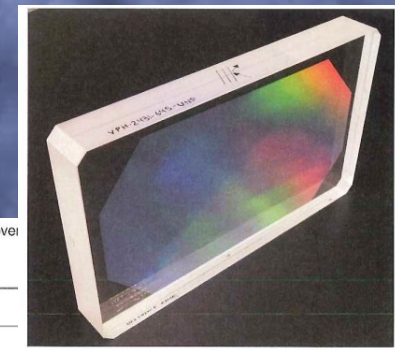
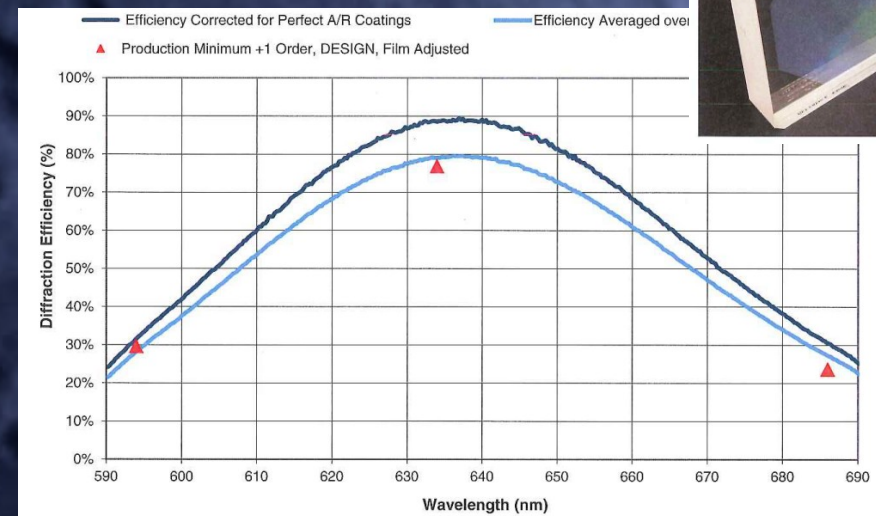
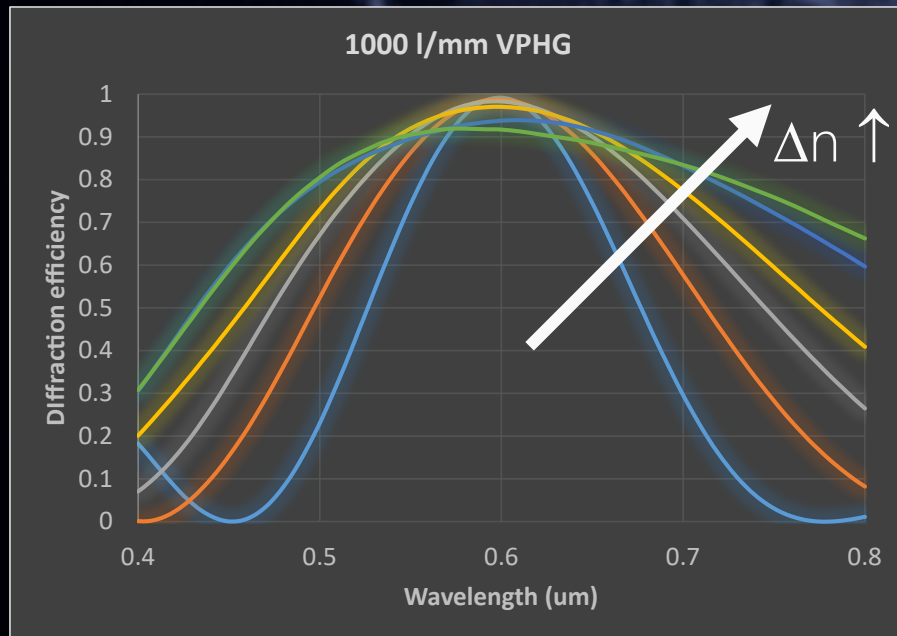
The diffraction occurs thanks to a periodic modulation of the refractive index in the volume of the material.

d and Δn are the key parameters that drive the diffraction efficiency

Periodic modulation of the refractive index Δn (usually sinusoidal).

Dispersing elements: VPHG

- The peak efficiency can be theoretically 100%. Easily, values of 90% are obtained;
- The devices are robust since the active material is usually embedded in between two glass windows and multilayer coatings can be applied;
- Large VPHGs can be produced if a big holographic setup is available;
- Large dispersion gratings with line density up to 6000 l/mm can be obtained;
- The device is easily customizable.



VPHG materials: DCG

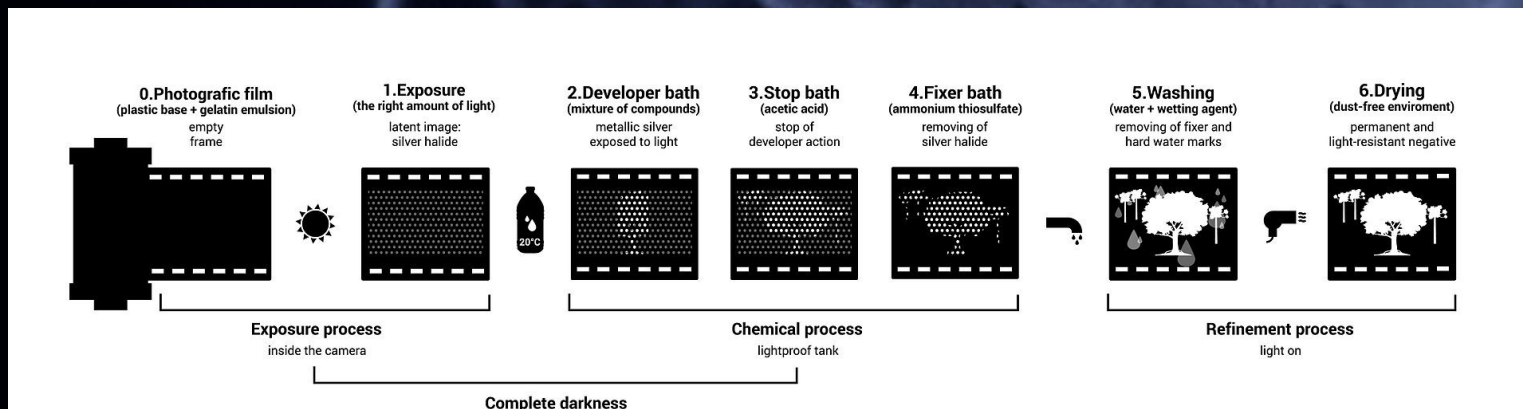
- Dichromated gelatin (DCG) has been studied for making phase holograms for more than 40 years. It contains Cr (toxic);
- It is surely **the best** material for holography: reference material for VPHG in astronomy;
- Modulation of the refractive index very large => Δn up to 0.15!
- Low sensitivity (high laser power);
- Transparent up to $2.5 \mu\text{m}$;
- Good homogeneity => very low scattering => high S/N.



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 - Good homogeneity => very low scattering => high S/N;
- BUT...**

A chemical development process is required (old camera films)



Microvoids are formed in the material



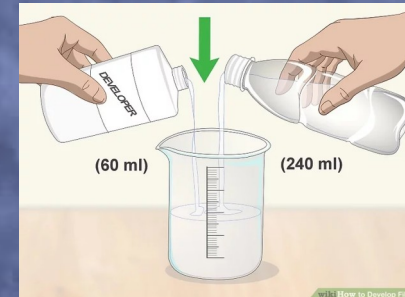
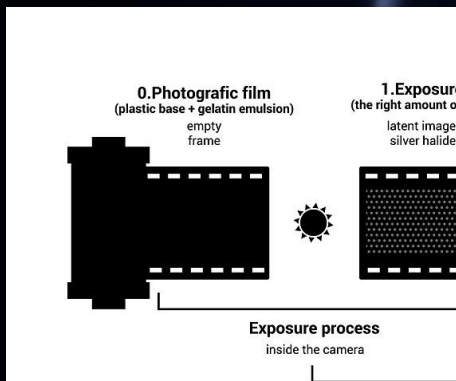
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- BUT...

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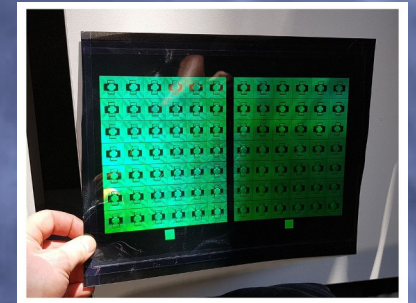
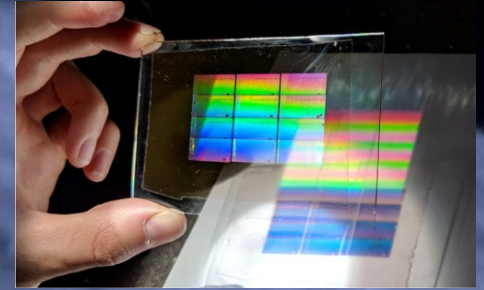


5!



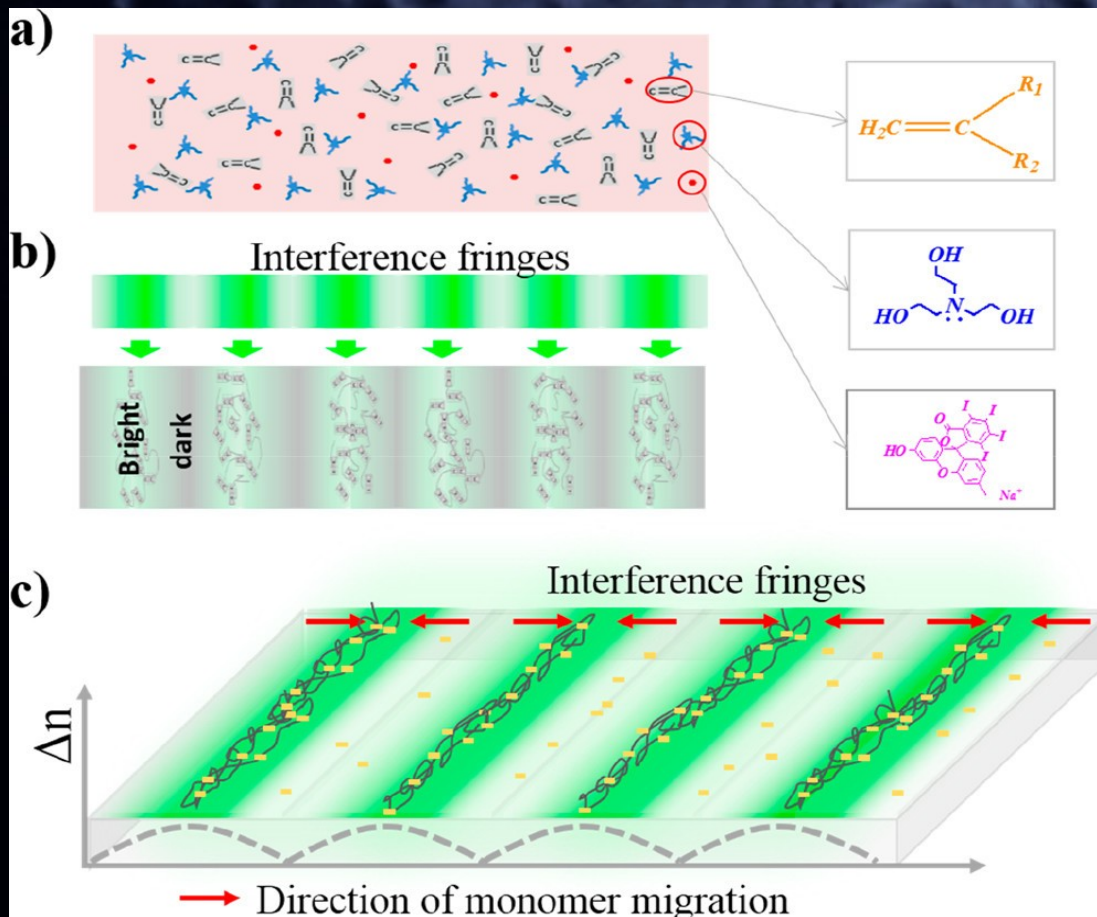
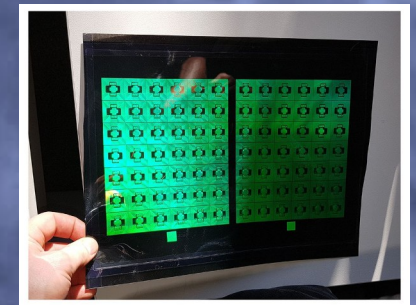
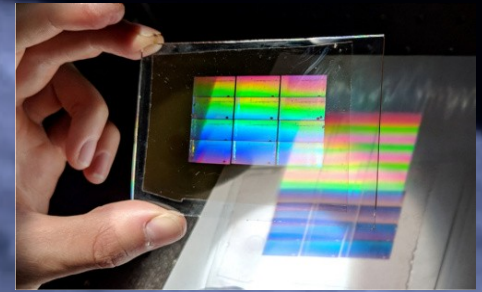
VPHG materials: Photopolymers

- High sensitivity (low laser intensity required);
- Easy control of the film thickness;
- Limited modulation ($\Delta n < 0.05$);
- Flexible, self-standing films;
- **Self-developing (no chemical post exposure process)!**



VPHG materials: Photopolymers

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We start from a mixture of monomer, photoinitiator, dispersed in a binder (a polymer)

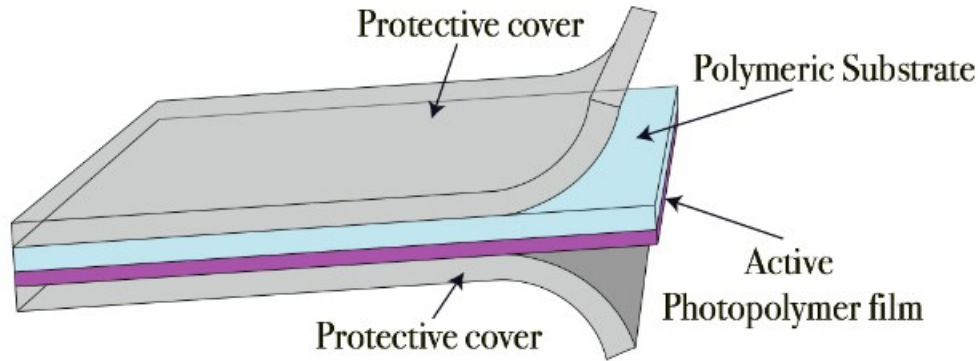
Migration of monomers from the dark to the bright areas;

Change in density = change in refractive index!

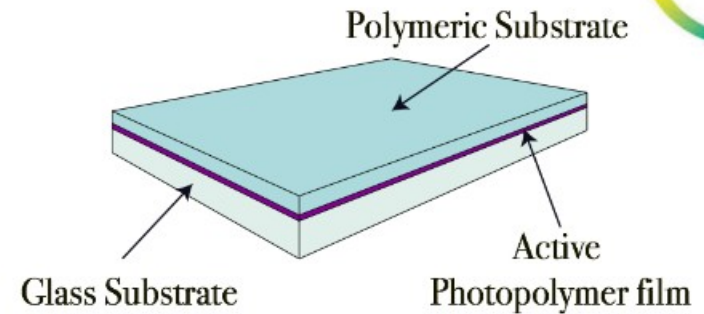
VPHG materials: Bayfol HX by Covestro

Bayfol® HX photopolymers by Covestro show a very simple and "easy to use" structure:

Film's Compositional Layers



Samples Scheme

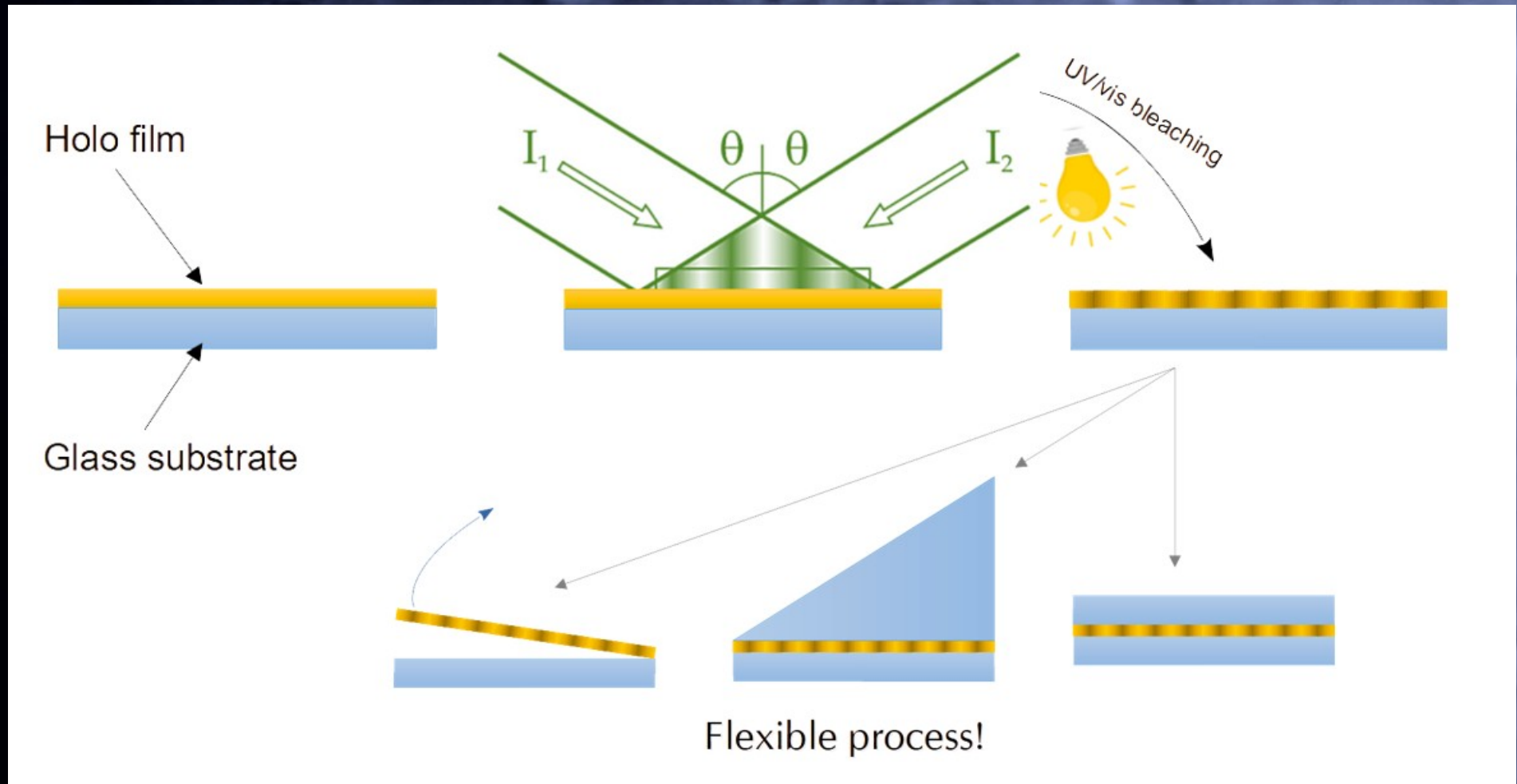


- The structure is like a protective layer of smartphone and tablet;
- Bayfol®HX photopolymers are available in different thicknesses and size (up to meter!).



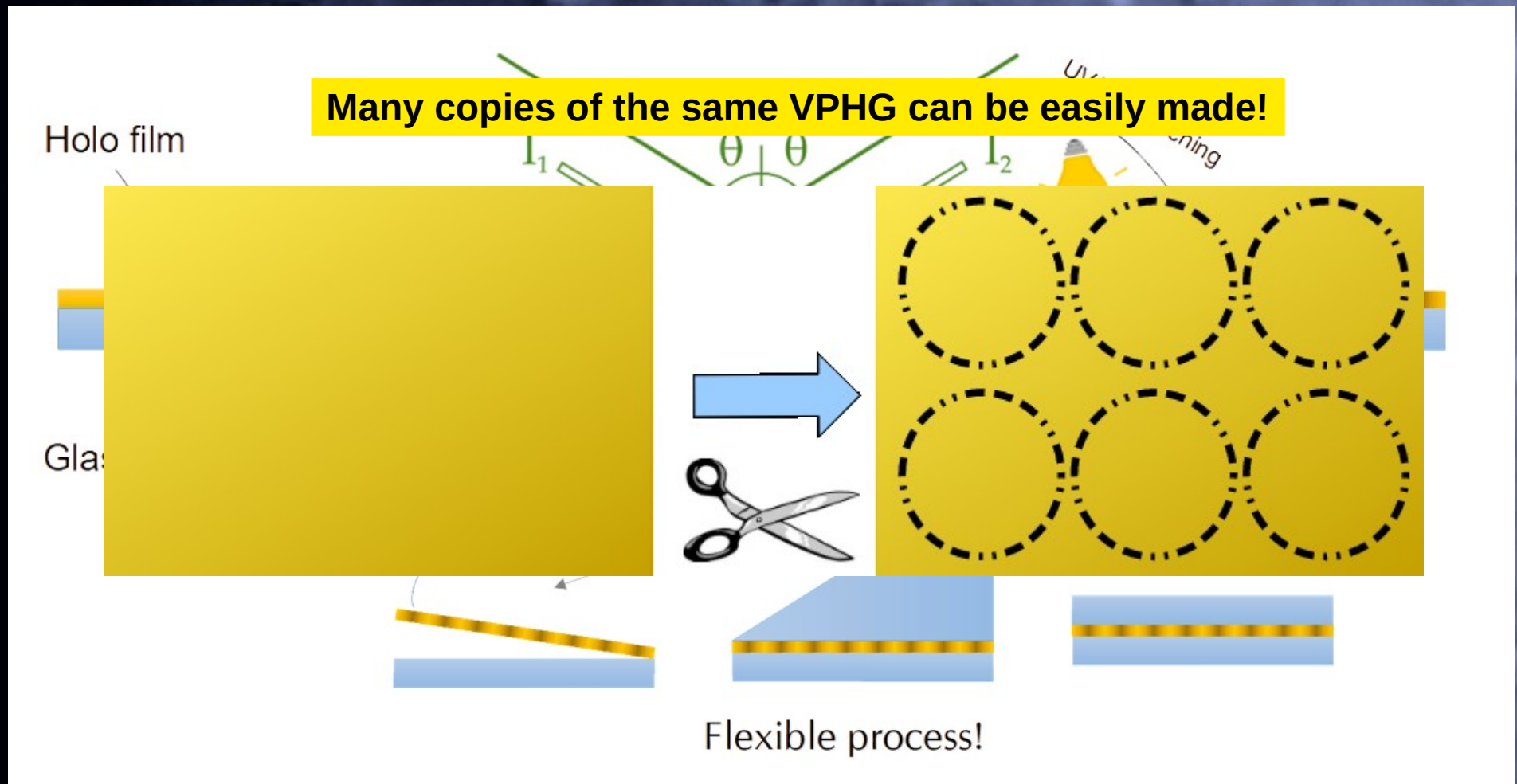
VPHG materials: Bayfol HX by Covestro

The process is highly flexible, allowing for a better optimization of the device. Once the element is written can be removed, coupled with prism, substrates, ..



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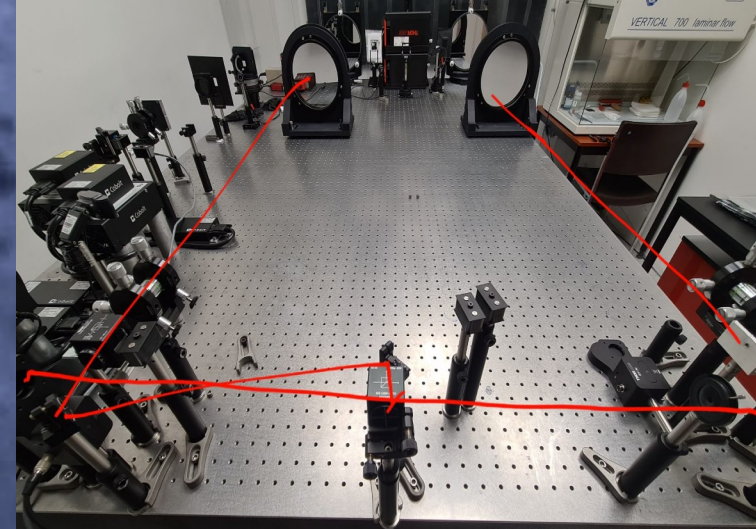
VPHG: what we can do...

Holographic set-up with RGB lasers;

VPHGs up to 190 mm in diameter working in the VIS and NIR also at high dispersion;

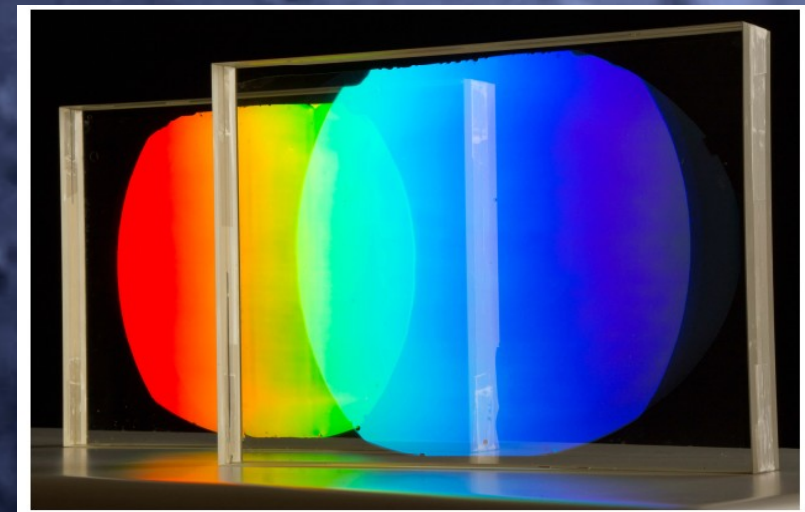
Non conventional VPHGs (multiplexed, multiorders) to increase spectral range and resolution;

More than 10 VPHG *made in OABr* available on observing facilities.



Bayfol® HX holographic films star in INAF telescope application

The Italian National Institute for Astrophysics (INAF) needed a highly efficient, diffractive light-guiding material to improve the performance of its specialized telescopes. Bayfol® HX holographic films won the starring role. These films are versatile, easily processed and offer great optical quality.

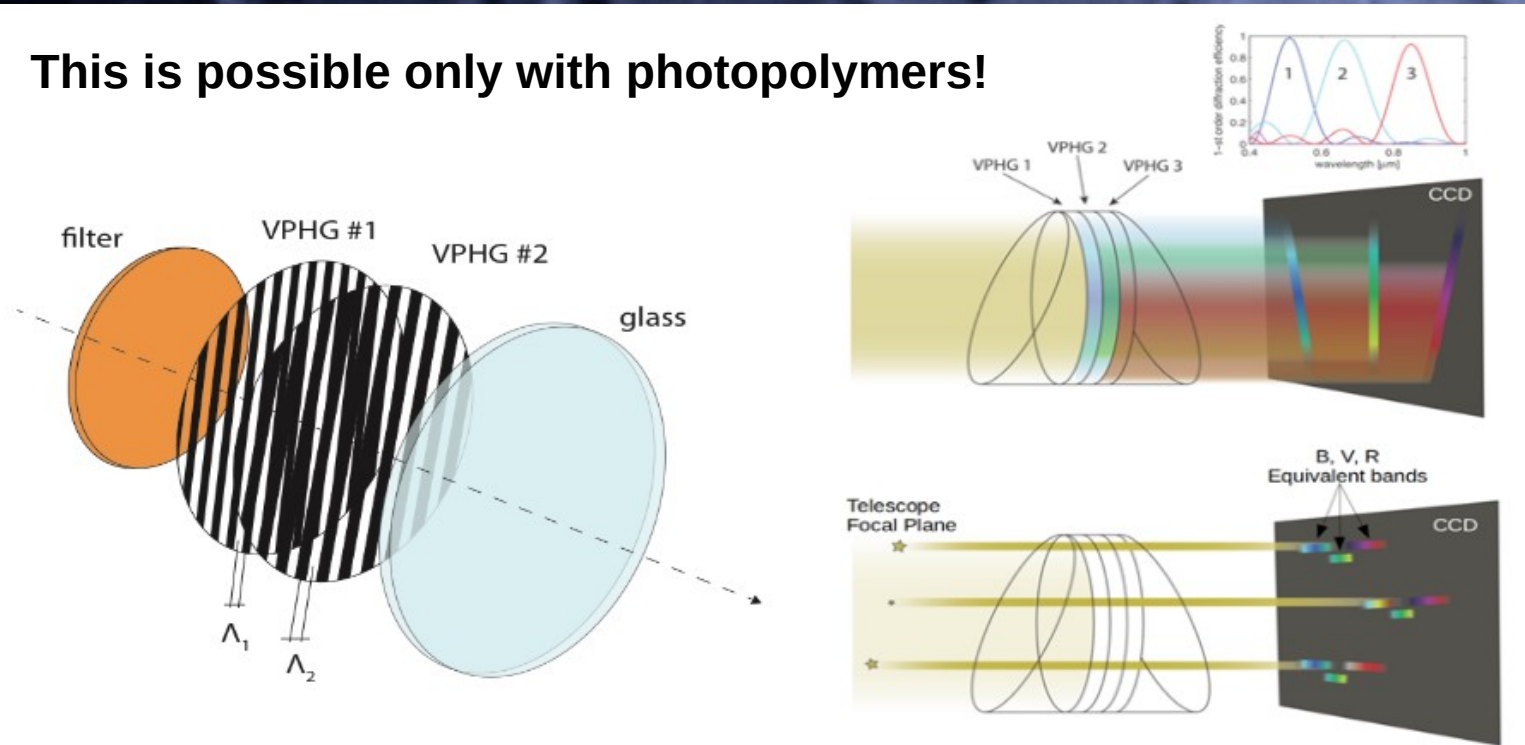


Example: Multiplexed VPHGs

ISSUE: increasing the resolution, the spectral range decreases due to the limited size of the detector. Multi exposures are necessary to cover a wide spectral range with the desired R.

IDEA: Combine more than one VPHGs with suitable clock angle in the same device to fill the detector in a similar fashion of an echelle. Target $R = 3000 - 5000$. Provide "high" res spectra only of some target features.

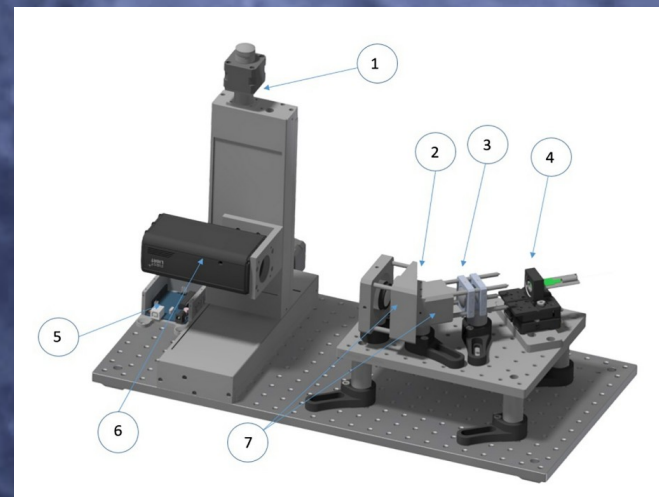
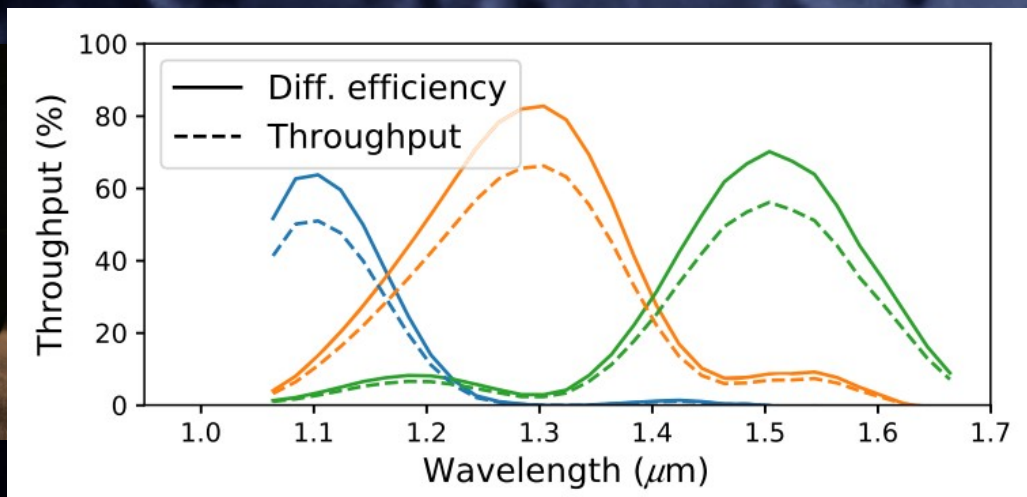
This is possible only with photopolymers!



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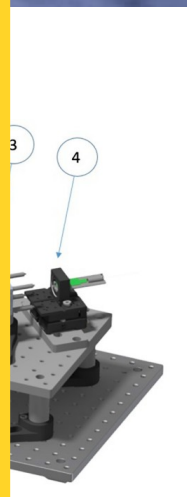
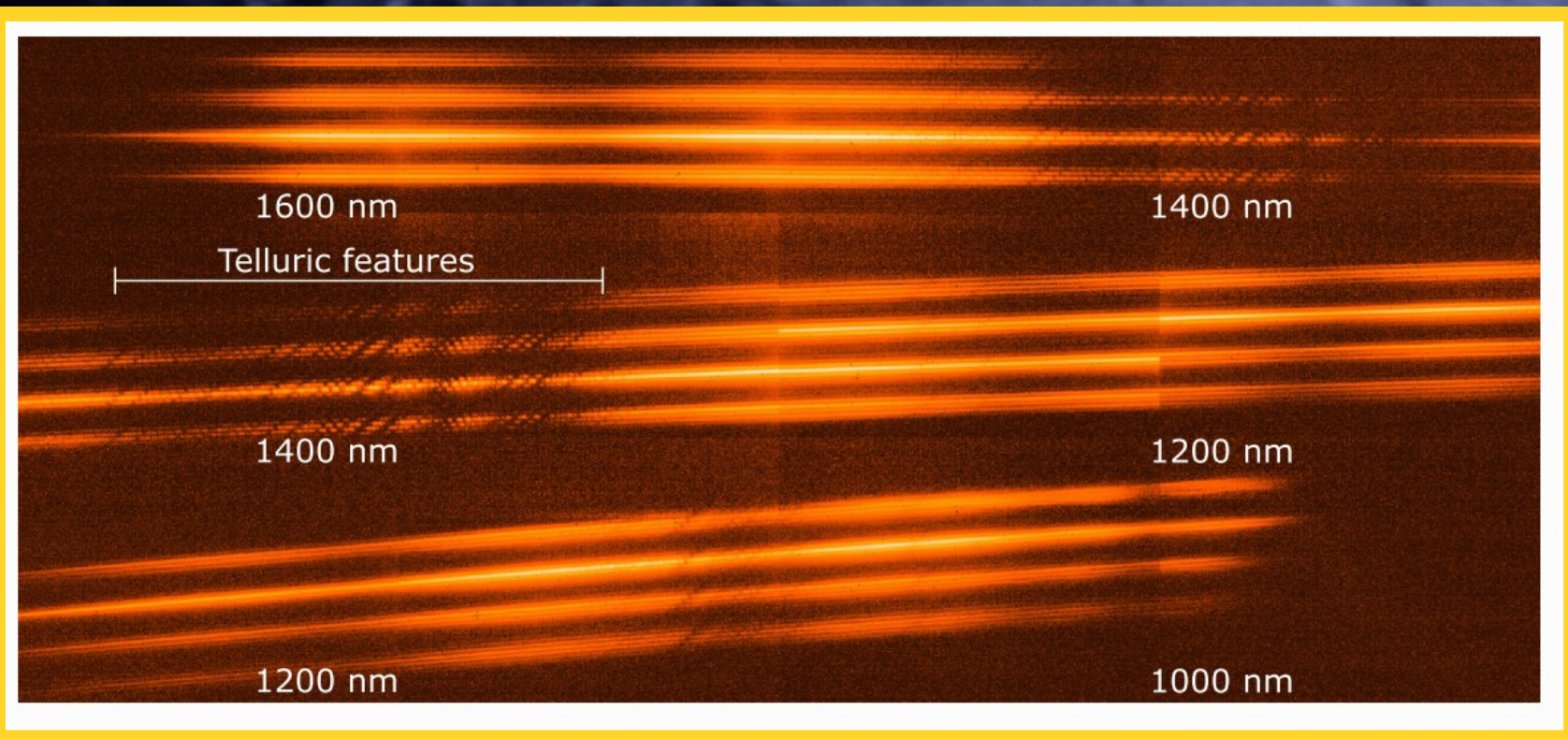


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Provided



Conclusions

We will experience an increase in the complexity, of both telescopes and their instrumentation;

New materials and processes can have a key role in this task;

The context and approach is multidisciplinary...not always easy;

Photochromic materials are a good example in this sense;

The optimization of the device requires the understanding of the materials at different levels;

Understand when to start such activities and how to support them.